

JINR



JOINT INSTITUTE FOR NUCLEAR RESEARCH



18 member states:

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6 associated members:

Egypt, Germany, Hungary, Italy, Serbia and Republic of South Africa

Laboratories:

High Energy Physics
Nuclear Problems
Theoretical Physics
Neutron Physics
Nuclear Reactions
Information Technologies
Radiation Biology
University Centre

Flerov Laboratory of Nuclear Reactions

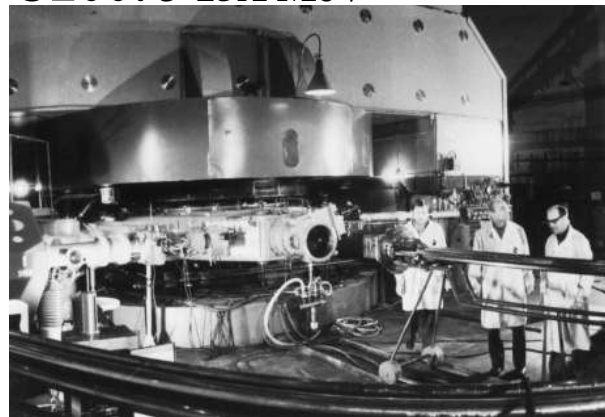
U400: 1-20A MeV



U400M: 6-100A MeV



U200: 3-15A MeV



IC-100: 0.5-1.2A MeV



MT-25 (electron accelerator)



fundamental physics

applied research

Main activity:

- **synthesis of SHE**
- **fusion-fission dynamics & multi-nucleon transfer**
- **RIB:**
 - **structure (${}^6,8,10\text{He}$, ${}^5,7\text{H}$,...)**
 - **reaction mechanisms**
- **nuclear theory**
- **applied research**
from polymeric membranes
to micro-chips for space satellites
- **education activity**

total staff ~ 400, scientists: > 100

Synthesis and study of heavy and superheavy neutron rich nuclei

- **Unexplored “north-east” area of the Nuclear Map**

- Nuclei below ^{208}Pb (astrophysical waiting point at $N=126$)
- Nuclei along the beta-stability line above Fermium
- Superheavy nuclei and the island of stability

- **Fusion reactions**

- What else can be done with fusion reactions ?
- Fusion reactions with RI beams ?

- **Multi-nucleon transfer reactions**

- Production of neutron-rich superheavy nuclei (U-like beam)
- Production of new neutron-rich heavy nuclei (Xe-like beam)
- Beams of accelerated fission fragments
- Experimental problems and separators of a new kind

- **Summary** (nearest and distant experiments)

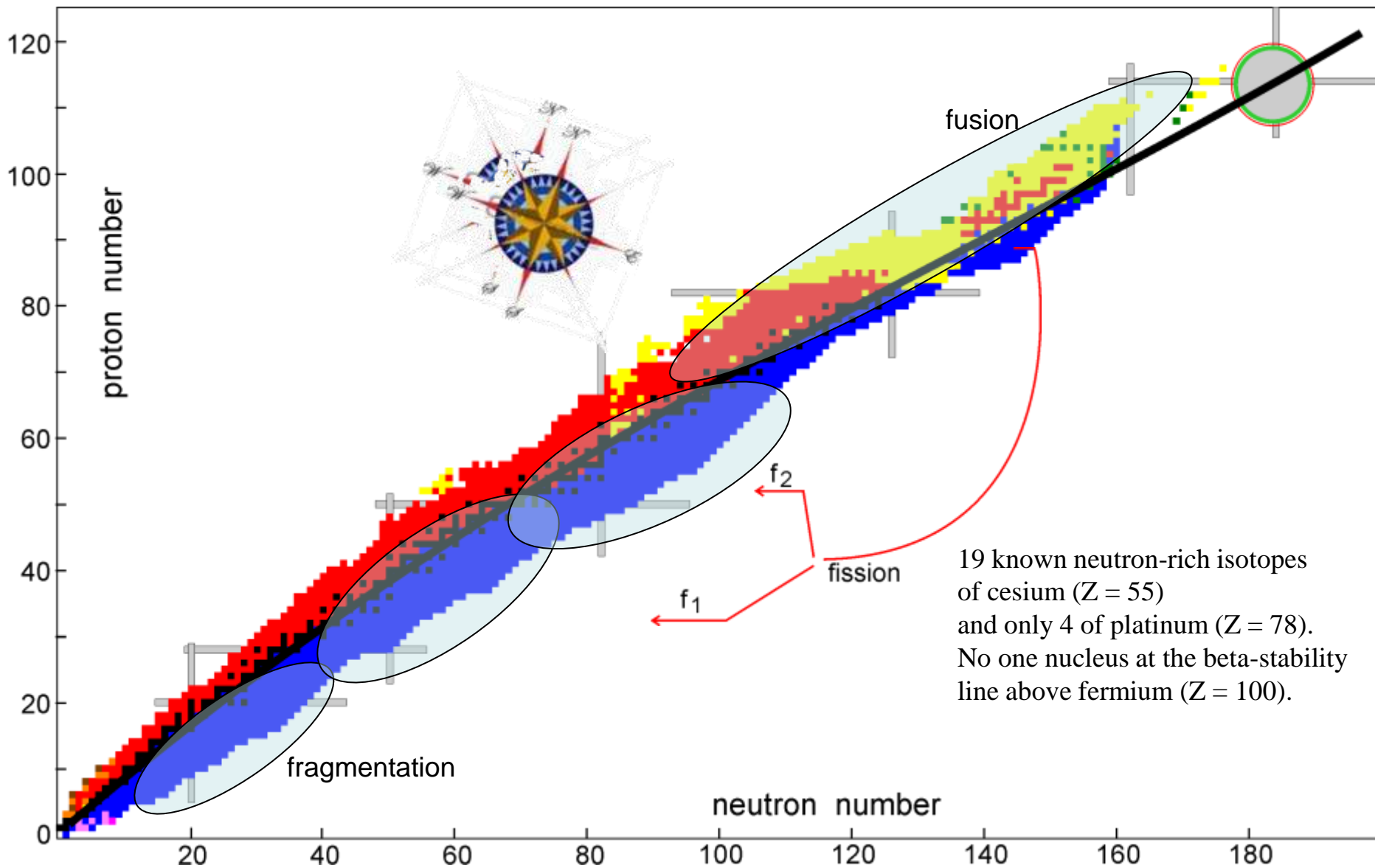


Valeriy Zagrebaev

Flerov Laboratory of Nuclear Reactions,
JINR, Dubna

for the 1-st International Symposium on the Science with KoRIA, 1-2 December, 2011

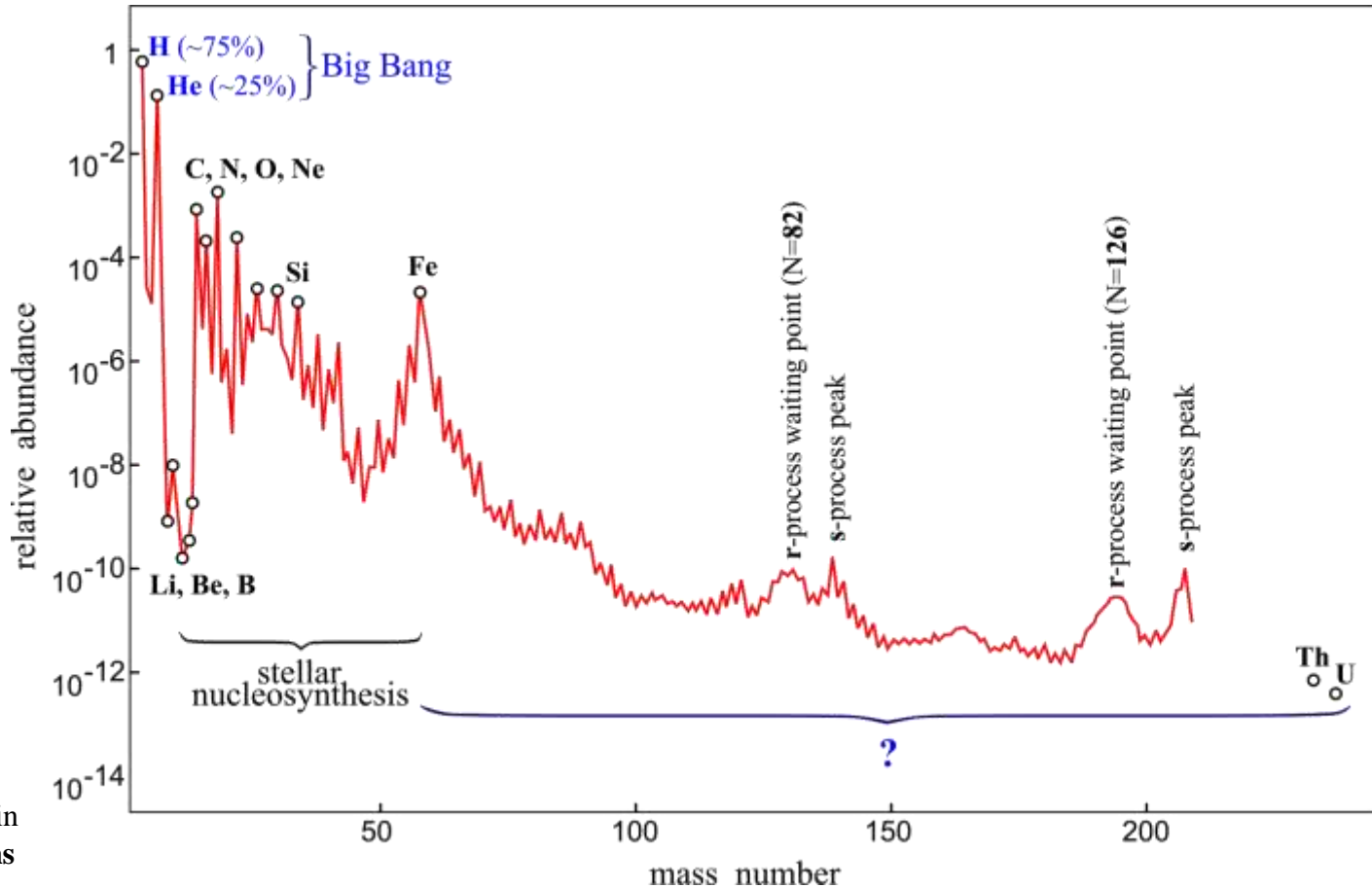
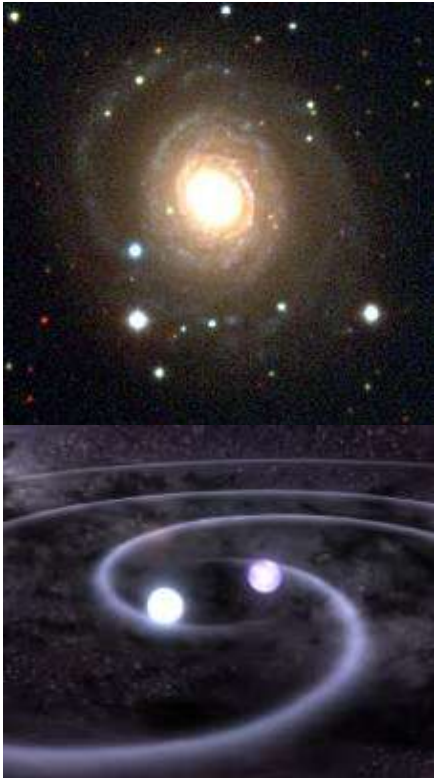
Unexplored “north-east” area of the nuclear map



Abundance of the elements in the Universe

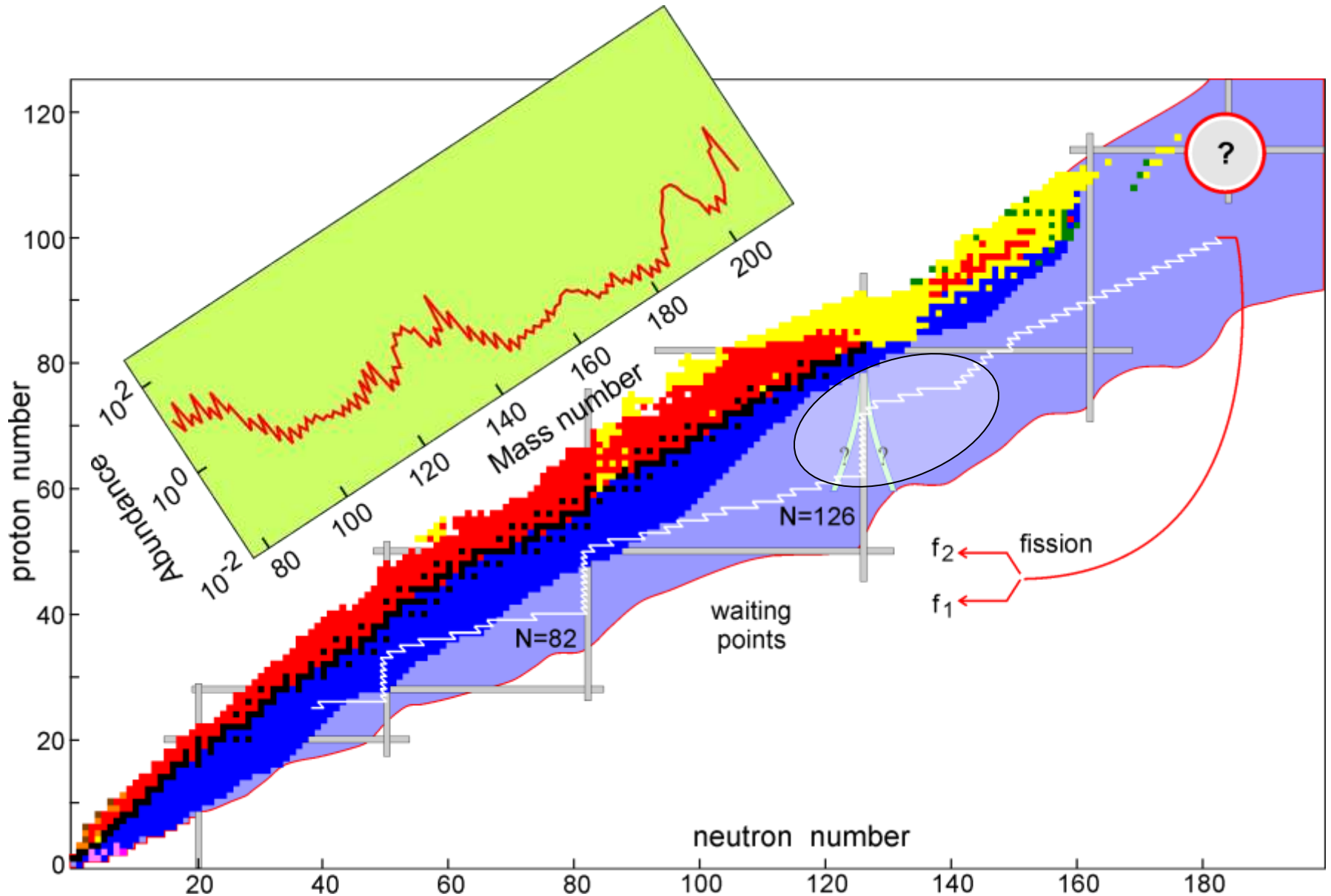
The 11 Greatest Unanswered Questions of Physics
(National Research Council, NAS, USA, 2002):

1. What is dark matter?
2. What is dark energy?
- 3. How were the heavy elements from iron to uranium made?**
4. Do neutrinos have mass?

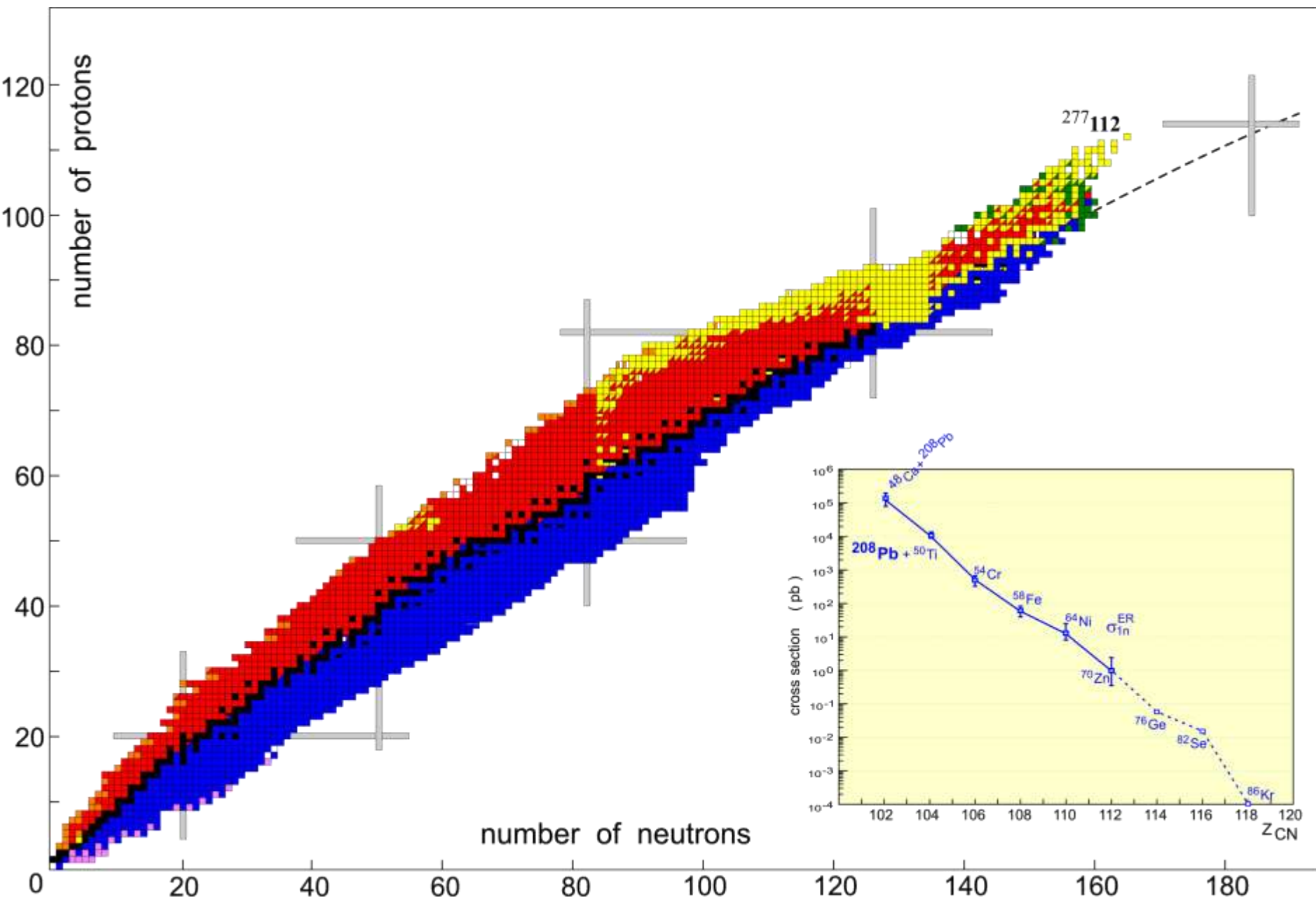


Strong neutron fluxes are expected in core-collapse supernova explosions or in the mergers of neutron stars.

r-process of nucleosynthesis and the neutron closed shell in the region of $N=126$



Nuclear Map in 2000



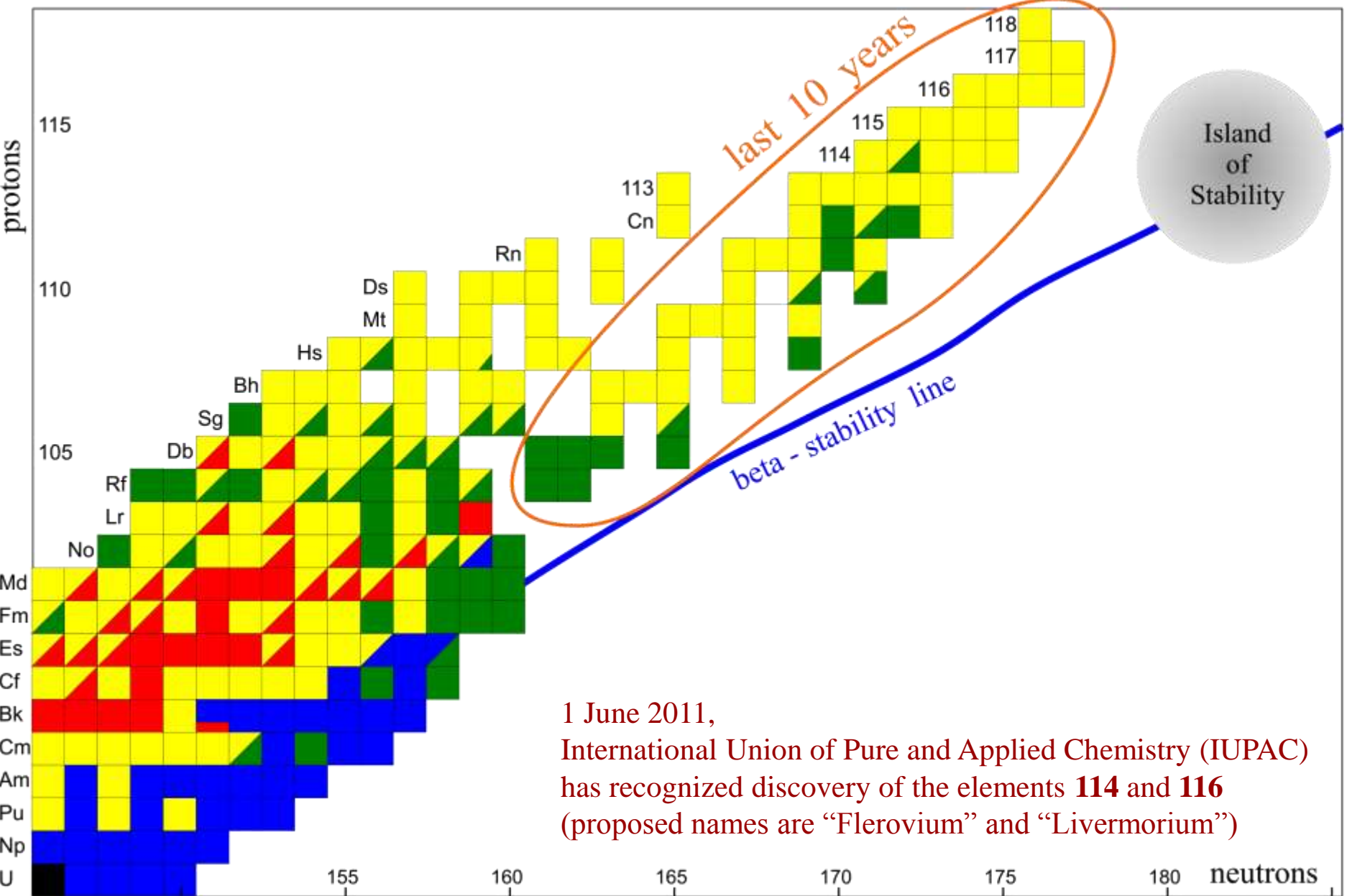
Synthesis of superheavy elements at FLNR (Dubna)

^{48}Ca induced fusion reactions with actinide targets



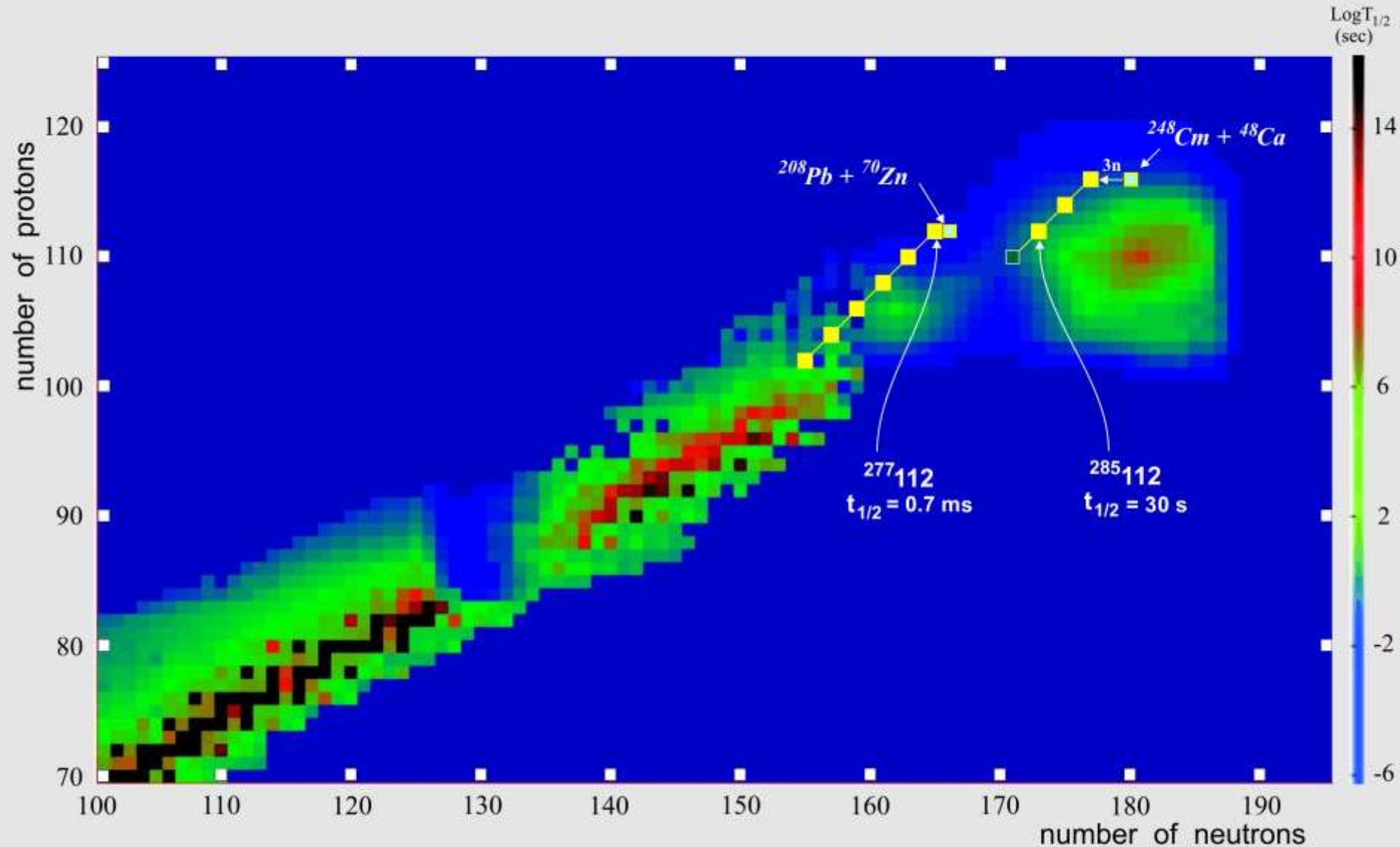
Yu. Oganessian, V. Utyonkov, et al. + Livermore + Oak Ridge

Great progress in synthesis of superheavy nuclei within last 10 years

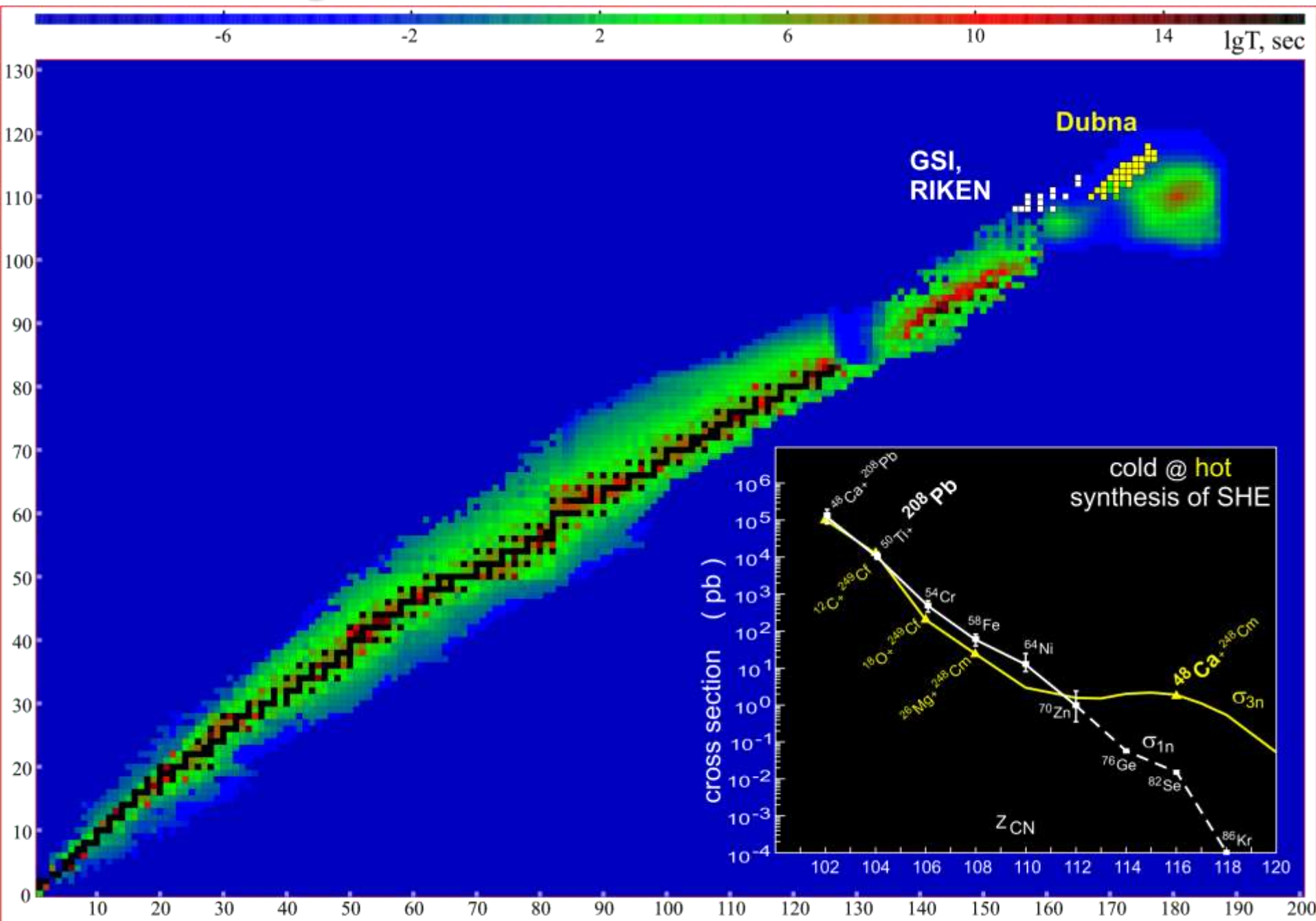


1 June 2011,
International Union of Pure and Applied Chemistry (IUPAC)
has recognized discovery of the elements **114** and **116**
(proposed names are "Flerovium" and "Livermorium")

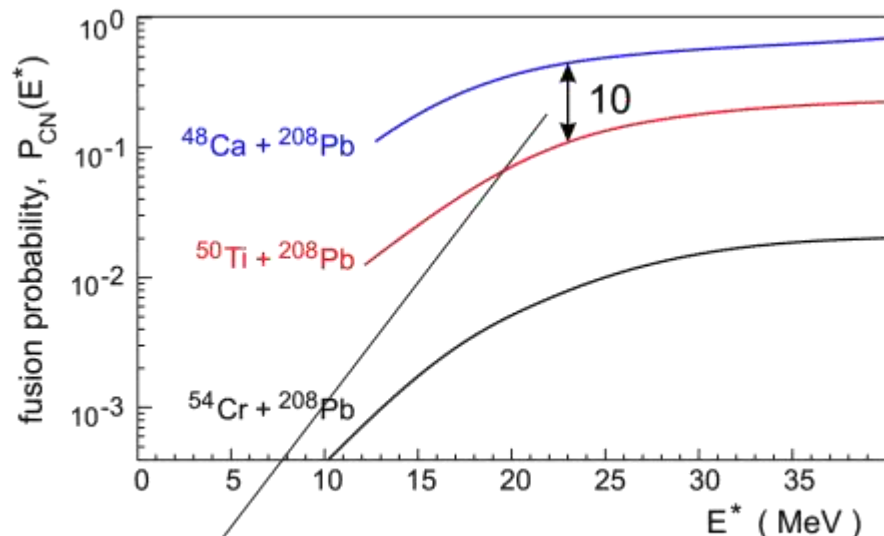
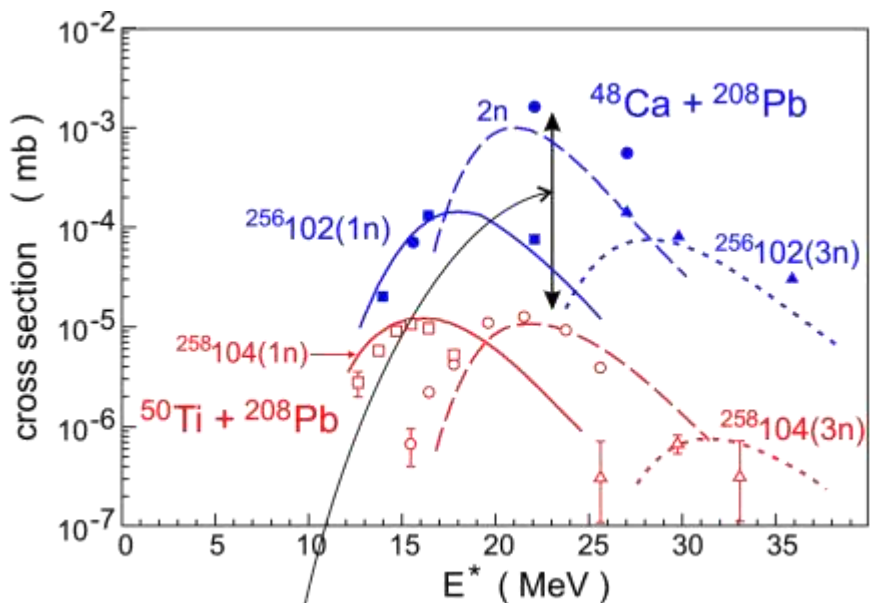
Approaching the Island of Stability



Drastic change in behavior of the cross sections (predictions of 2002)

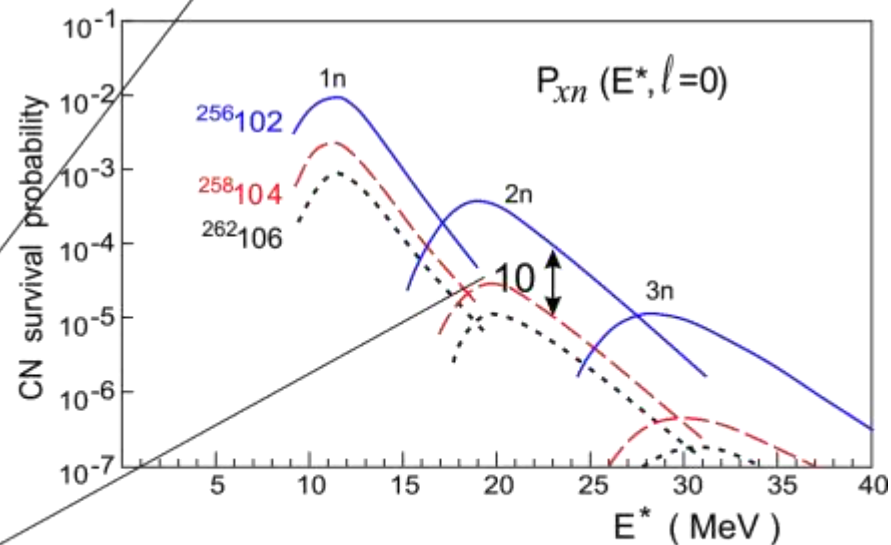


Epoch of ^{48}Ca is almost over. How much is ^{50}Ti worse?



$$\frac{\sigma(^{50}\text{Ti})}{\sigma(^{48}\text{Ca})} = \frac{1}{100}$$

	B_{LD}	δW	B_f	E_n
$^{256}_{102}$	1.26	4.48	5.7	7.1
$^{258}_{104}$	0.77	4.49	5.3	7.6



$$\sigma_{\text{EvR}}^{xn}(E) = \frac{\pi \hbar^2}{2\mu E} \sum_{\ell} (2\ell+1) P(\ell, E) \cdot P_{\text{CN}}(\ell, E^*) \cdot P_{xn}(\ell, E^*)$$

Beyond ^{48}Ca : ^{50}Ti and ^{54}Cr induced fusion reactions

Ti beam:

Cr beam:

TASCA

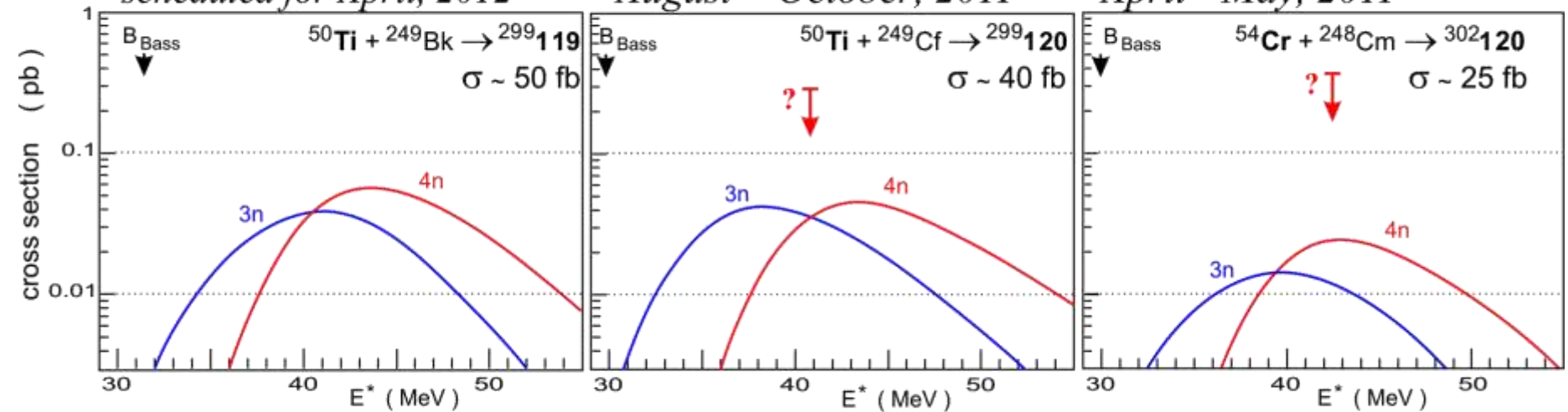
scheduled for April, 2012

TASCA

August - October, 2011

SHIP

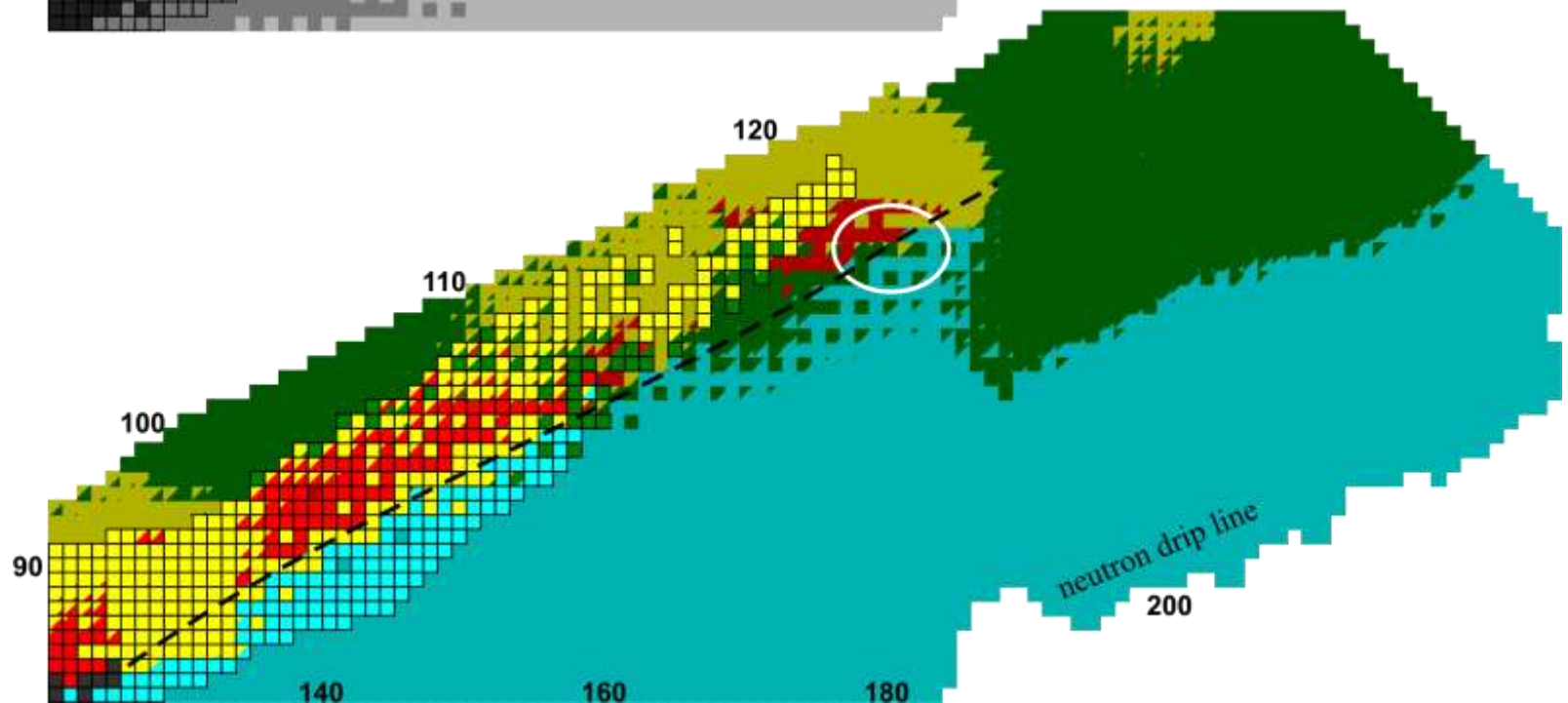
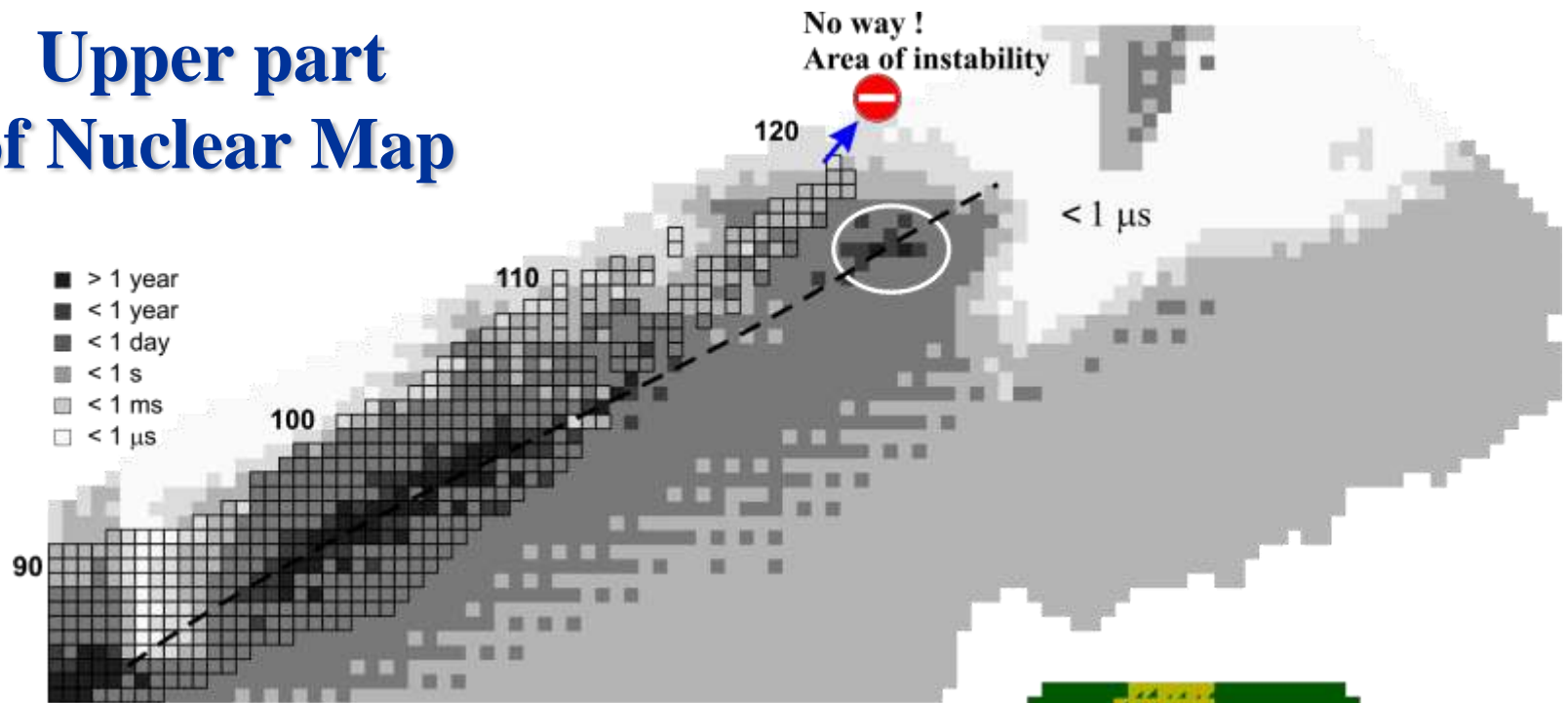
April - May, 2011



factor $\frac{1}{20}$ as compared to ^{48}Ca

*Maybe these elements are the last ones
which will be synthesized in nearest future !?*

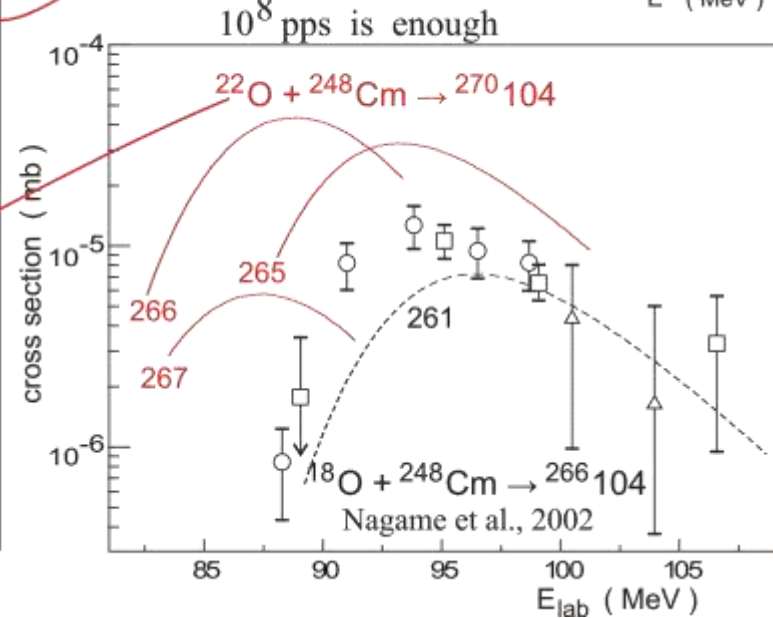
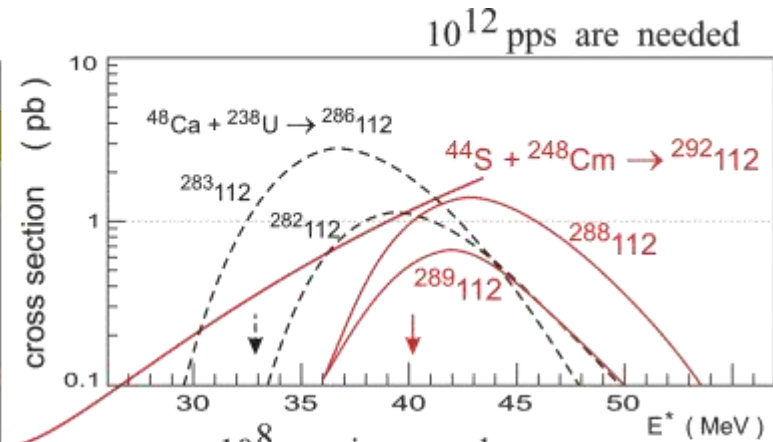
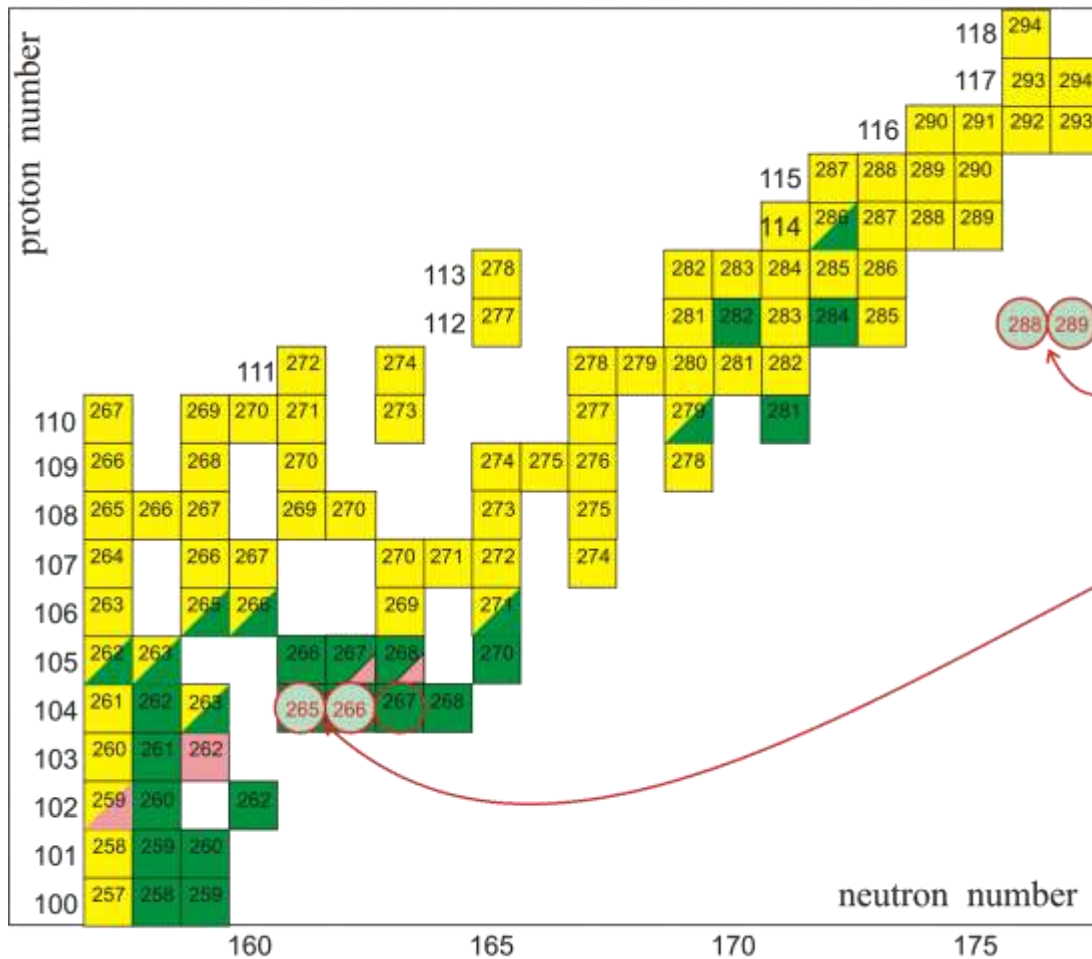
Upper part of Nuclear Map



How can we produce new heavy and superheavy neutron-rich nuclei ?

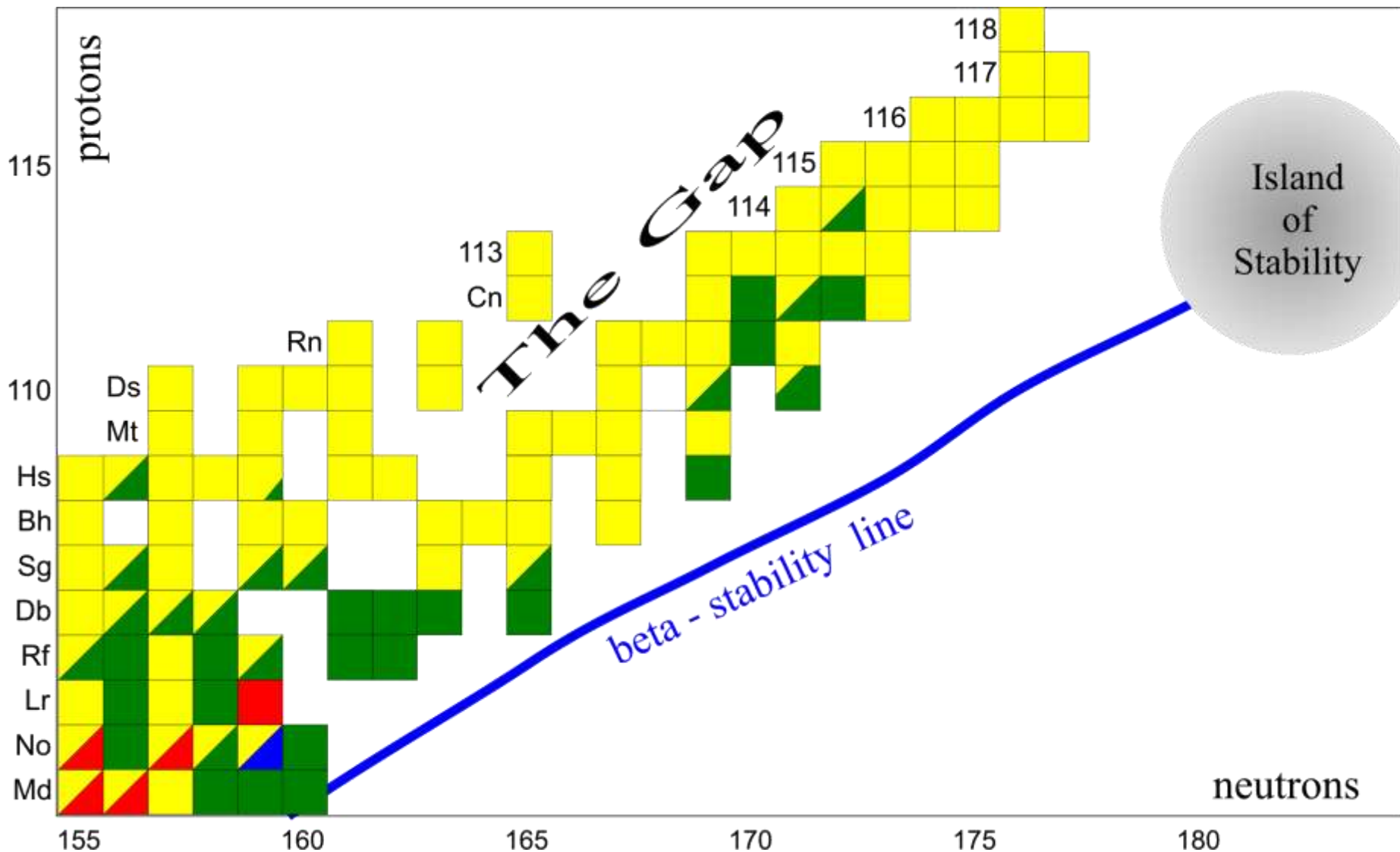
- 1. Fusion reactions of stable nuclei**
- 2. Fusion reactions with radioactive beams ?**
(e.g., $^{22}\text{O} + ^{248}\text{Cm}$, ...)
- 3. Multi-nucleon transfer reactions**
- 4. Neutron capture processes** (pulsed reactors of new generation)

Use of low-energy Radioactive Ion Beams for production of neutron rich superheavy nuclei ?

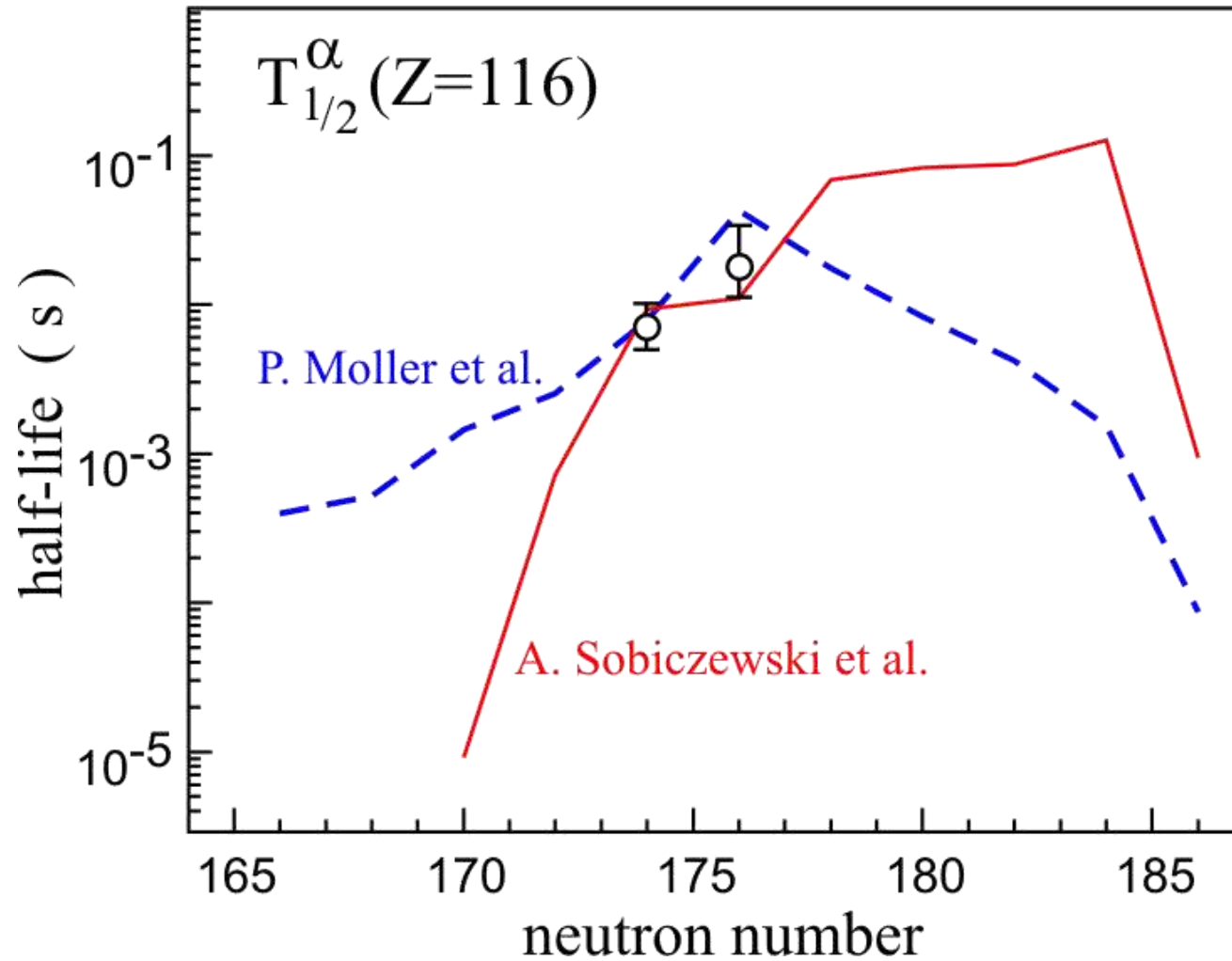


No chances today. But in future ?

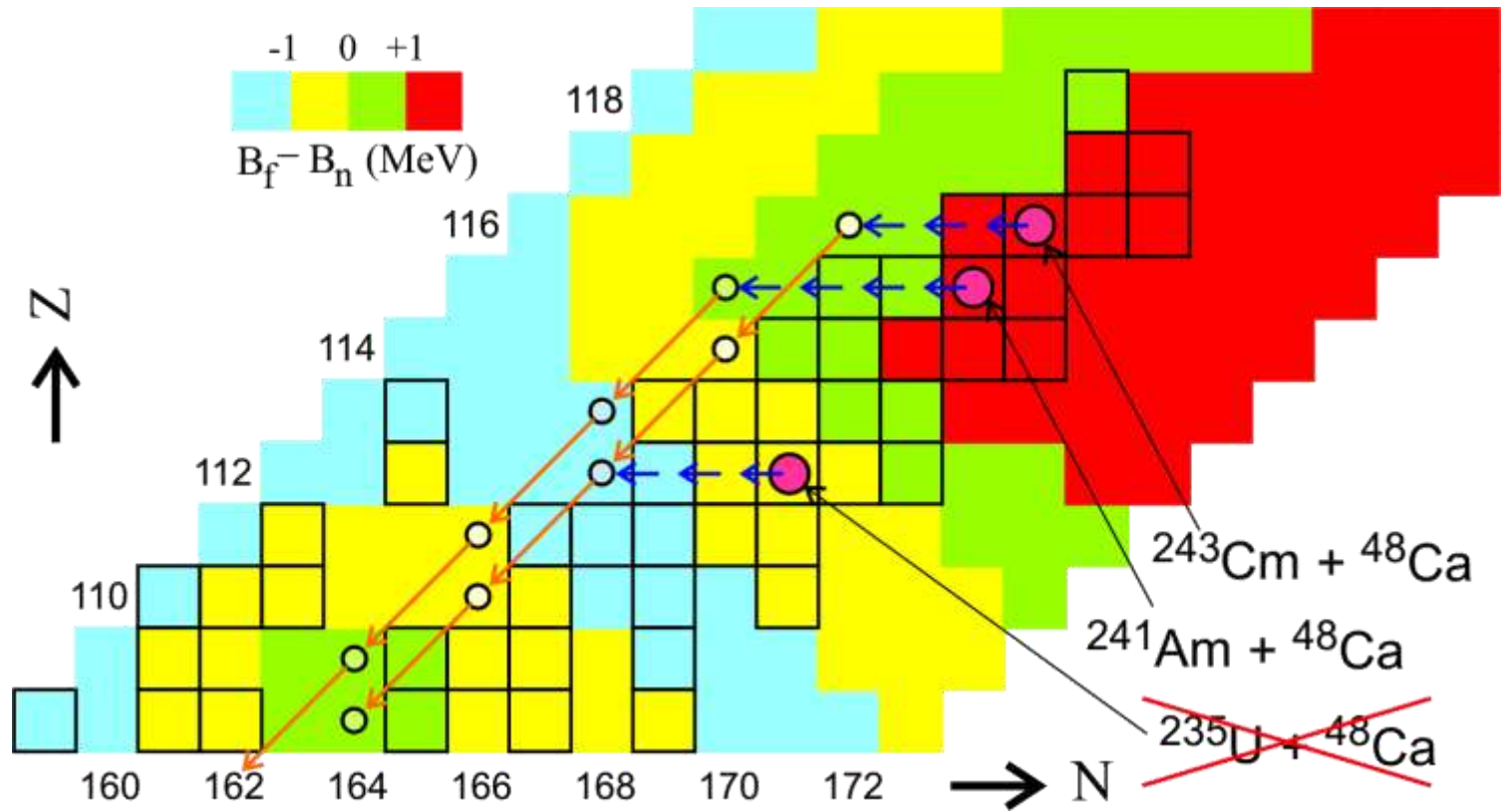
First, we should darn the gap in superheavy mass area?



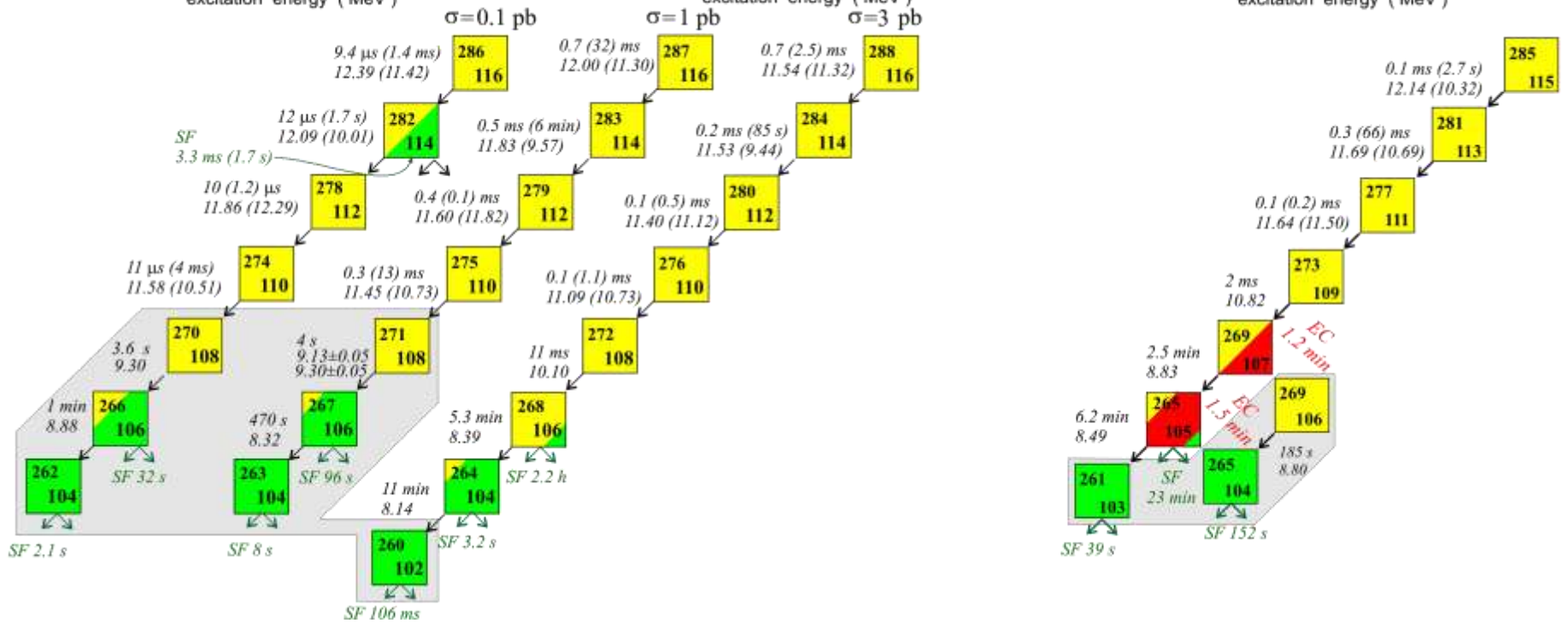
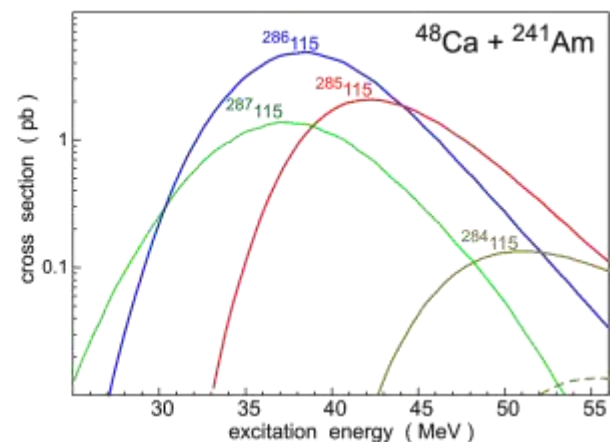
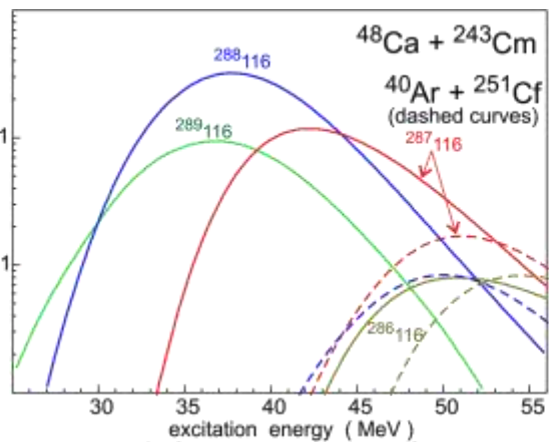
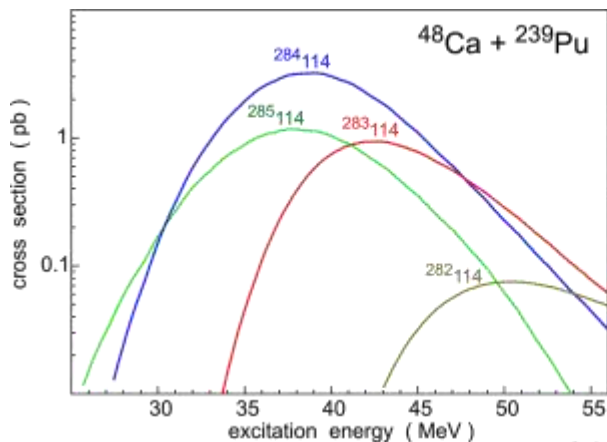
Our today ability of predictions in superheavy mass area



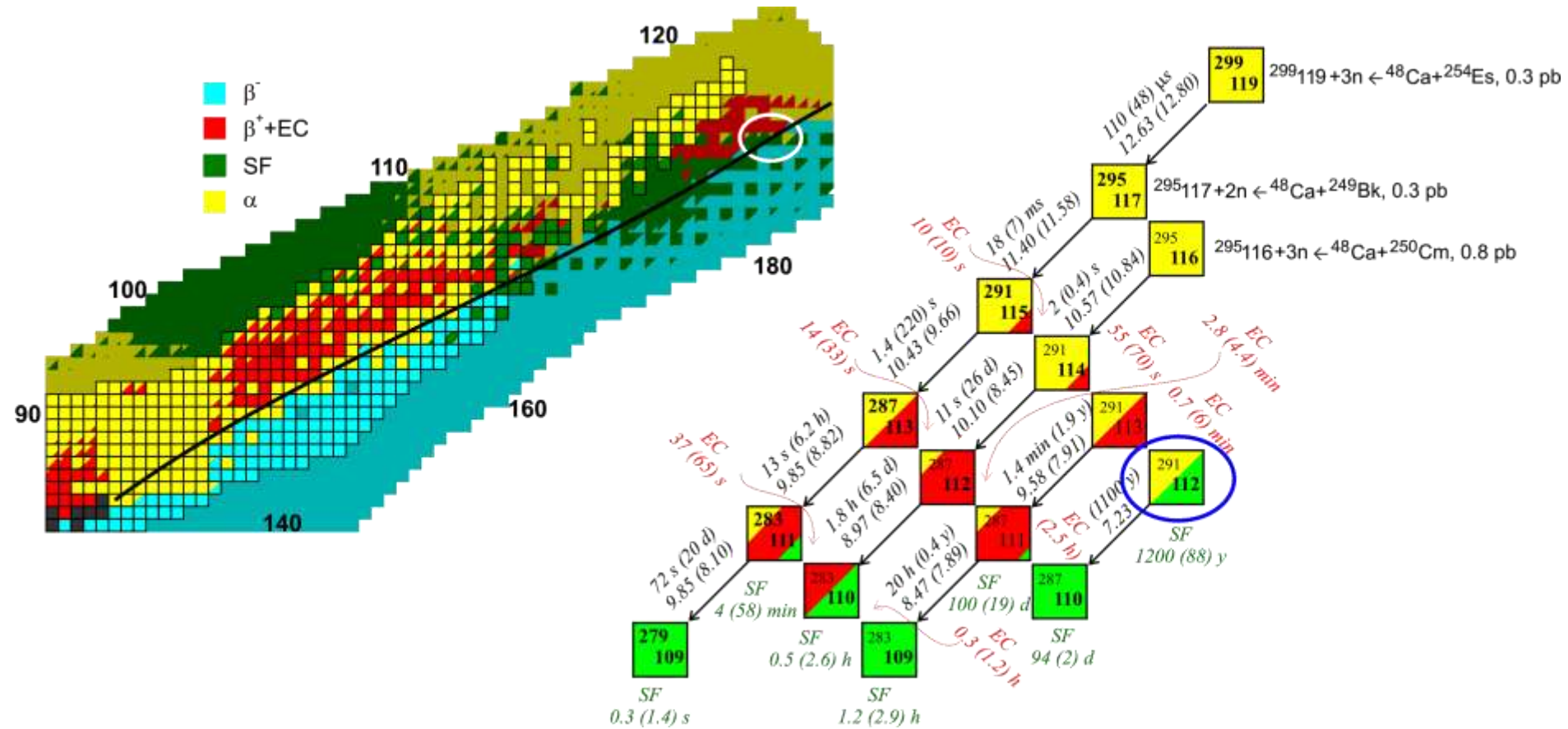
It is easier to darn the gap from above !



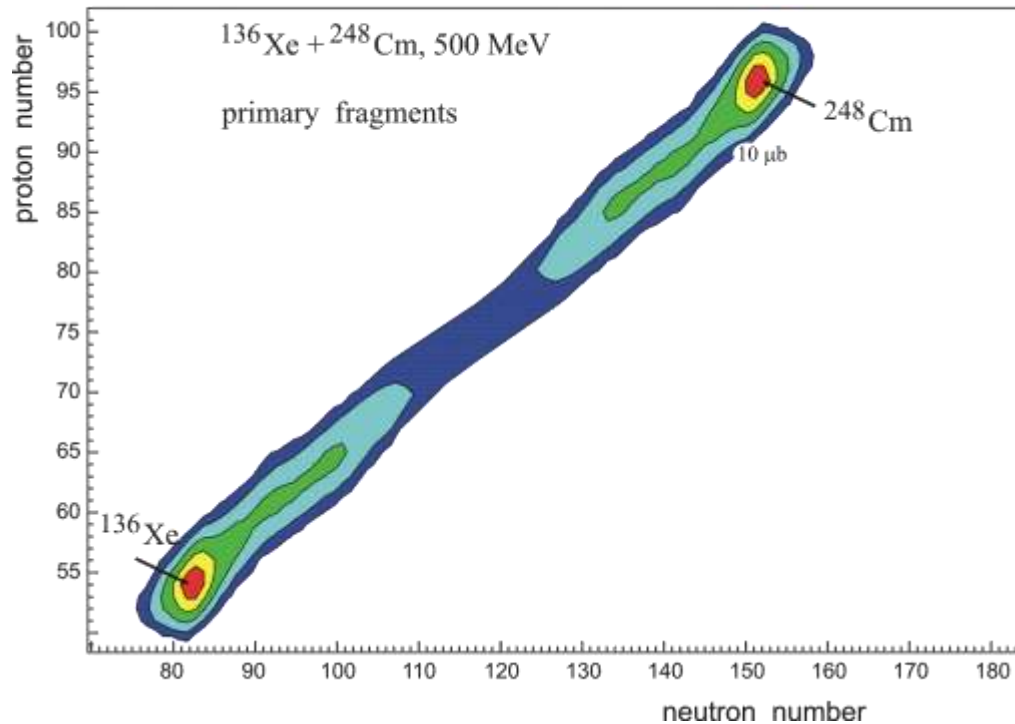
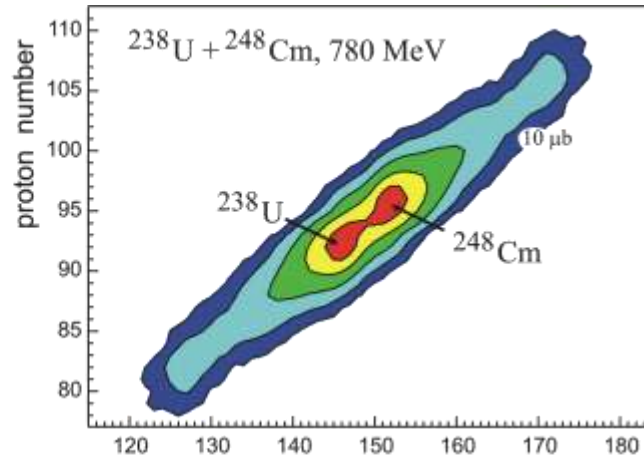
Cross sections are high enough to perform experiments at available facilities



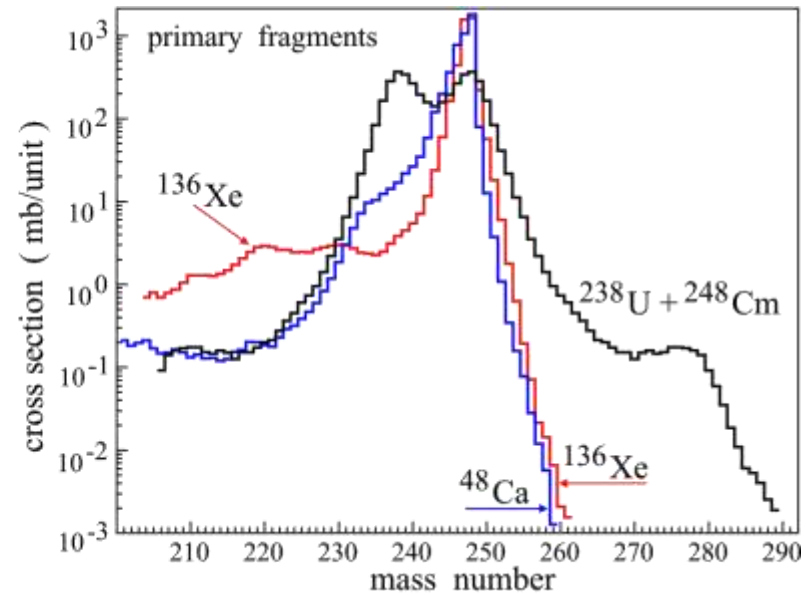
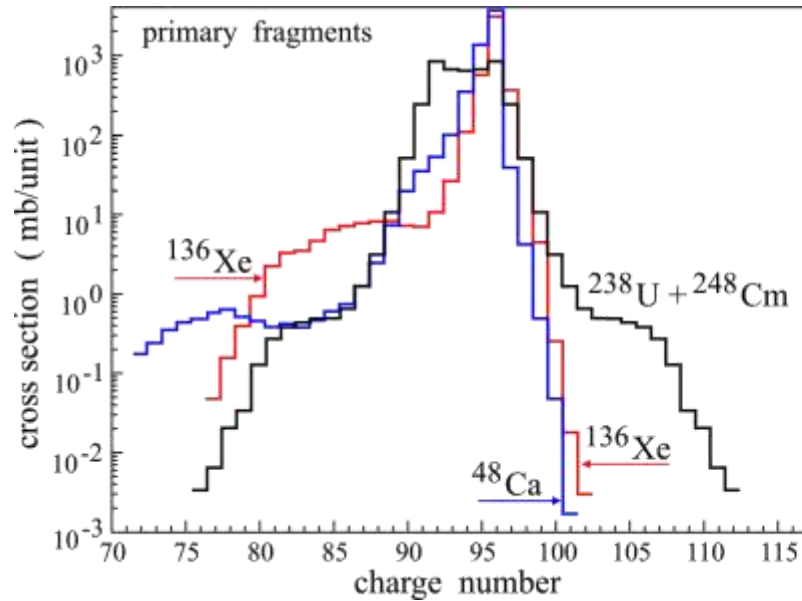
Narrow pathway to the island of stability !



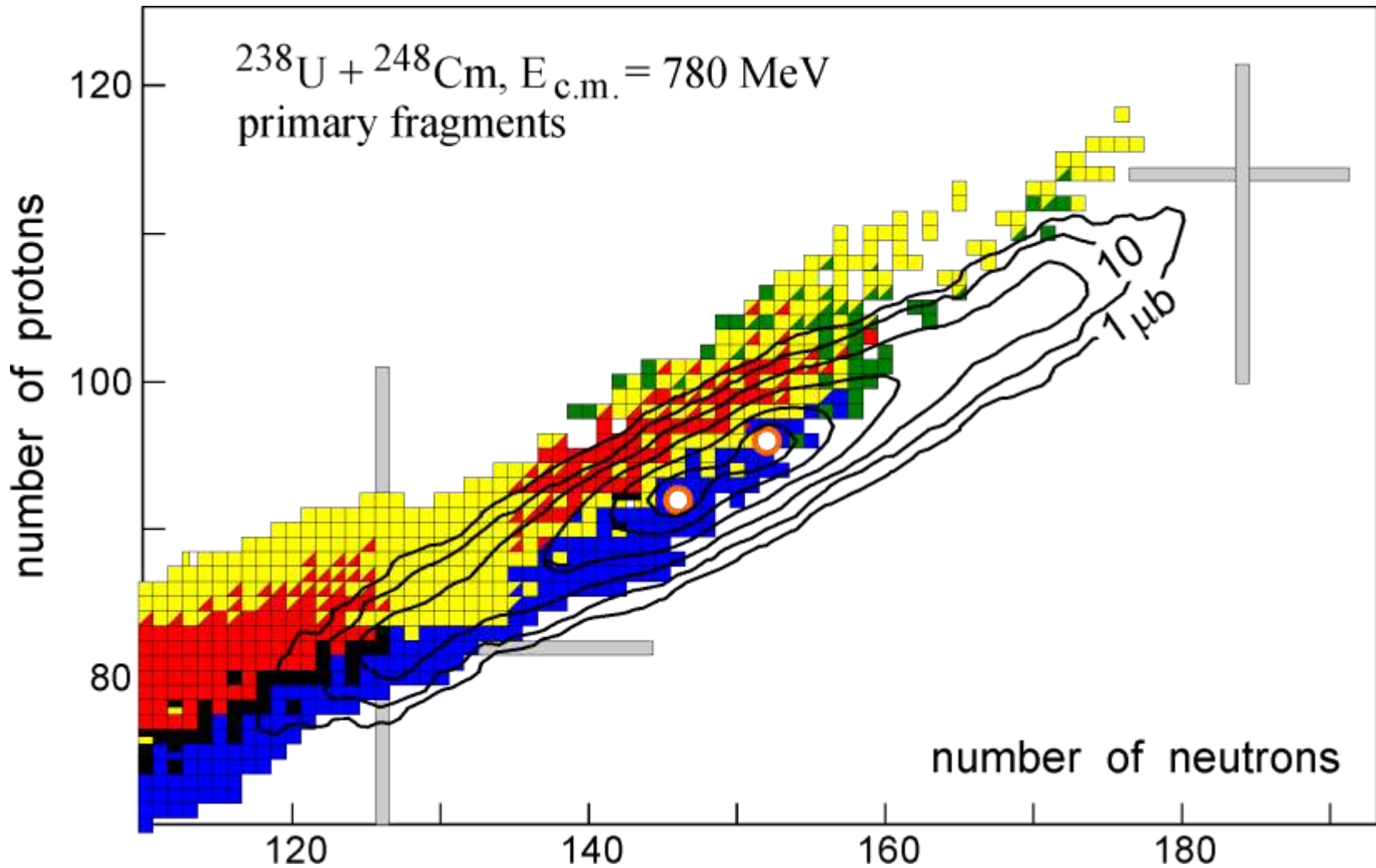
Production of SH nuclei in multi-nucleon transfer reactions (choice of reaction is very important)



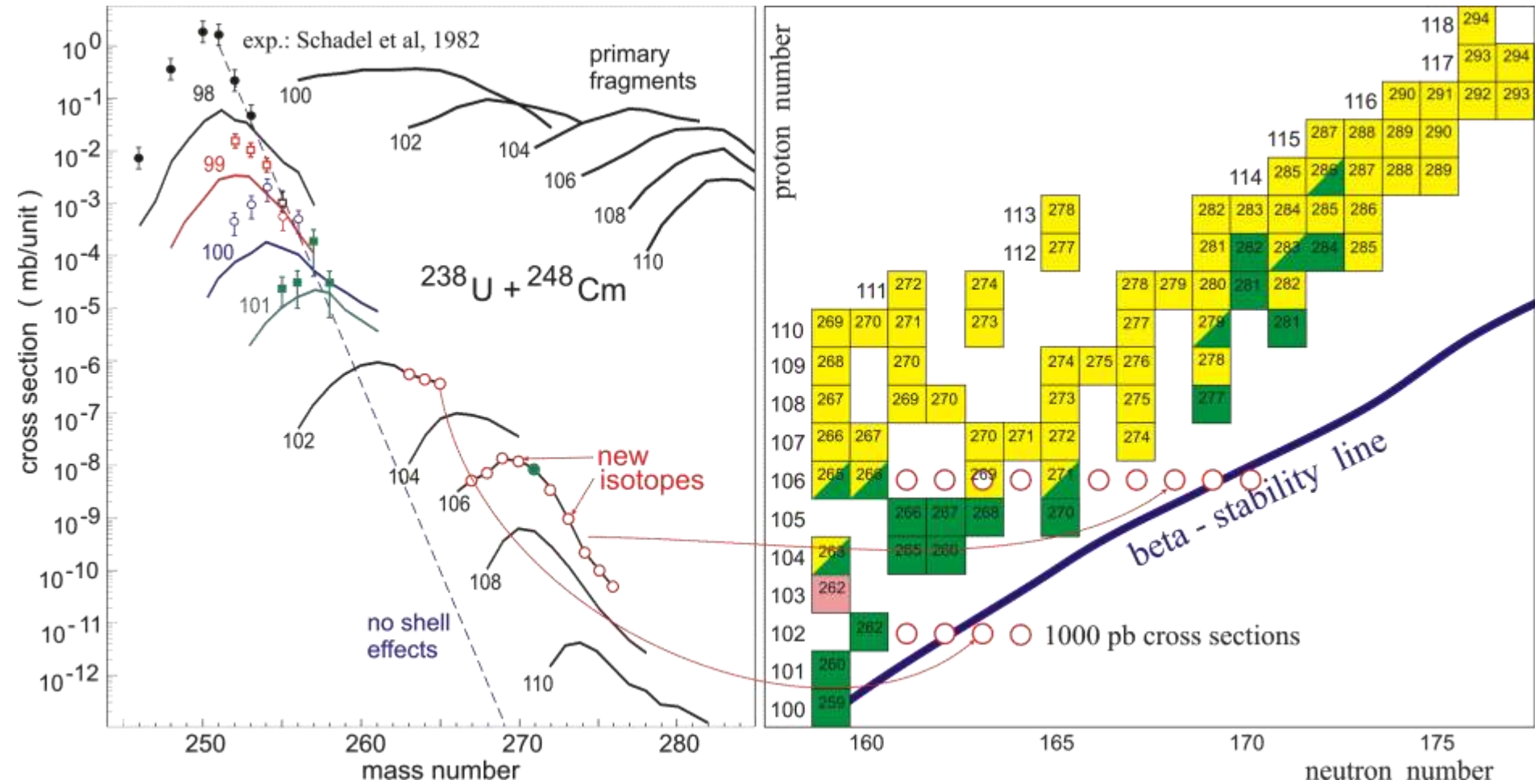
Only U-like beams give us a chance to produce neutron rich SH nuclei in transfer reactions



$^{238}\text{U} + ^{248}\text{Cm}$. Primary fragments

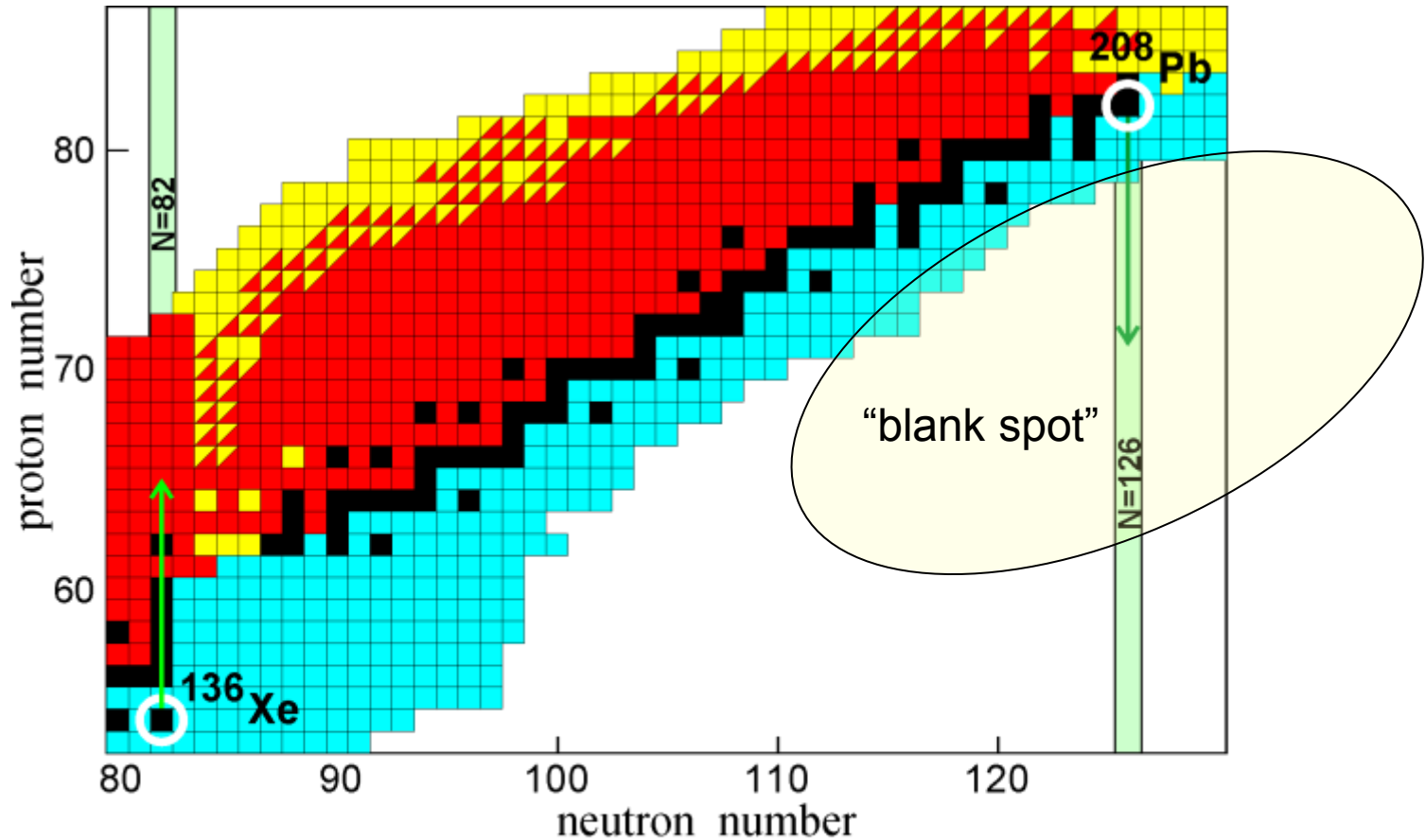


Study of transfermium nuclei along the line of stability becomes possible at last

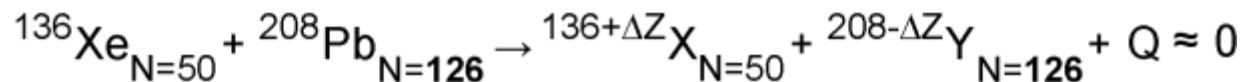


Production on NEW heavy nuclei in the region of N=126

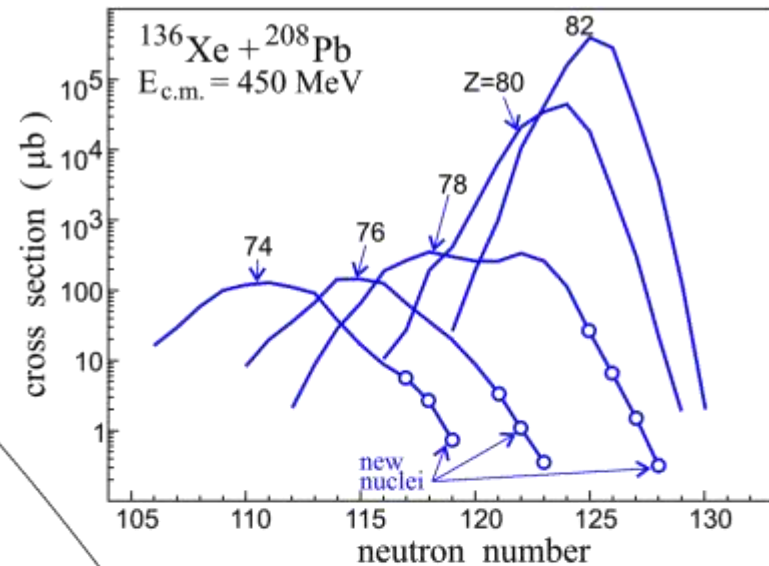
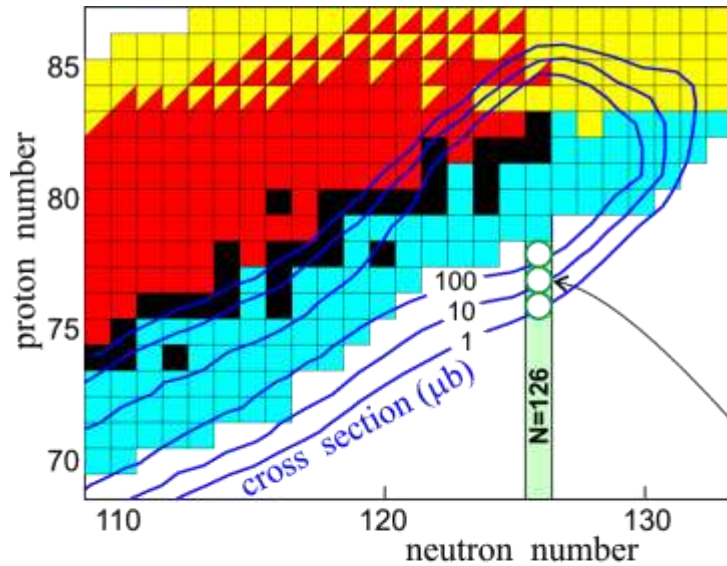
(Zagrebaev and Greiner, Phys. Rev. Lett., 2008)



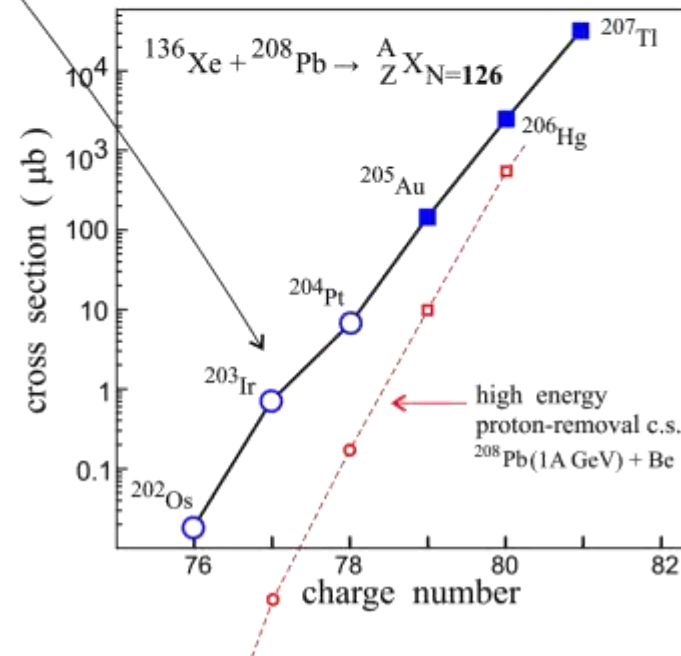
proton transfer along the neutron closed shells:



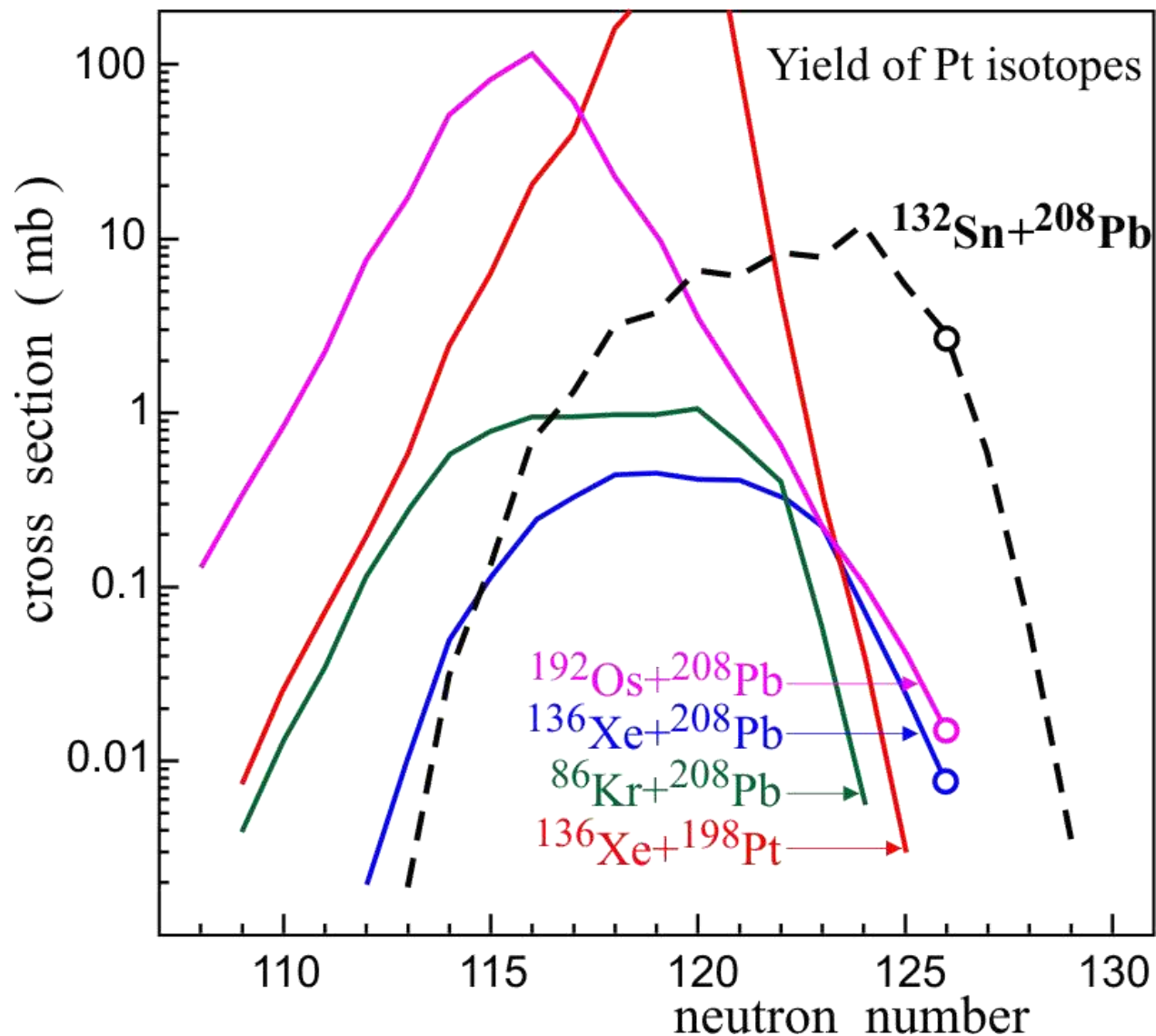
Production on new heavy nuclei in the $\text{Xe} + \text{Pb}$ collisions



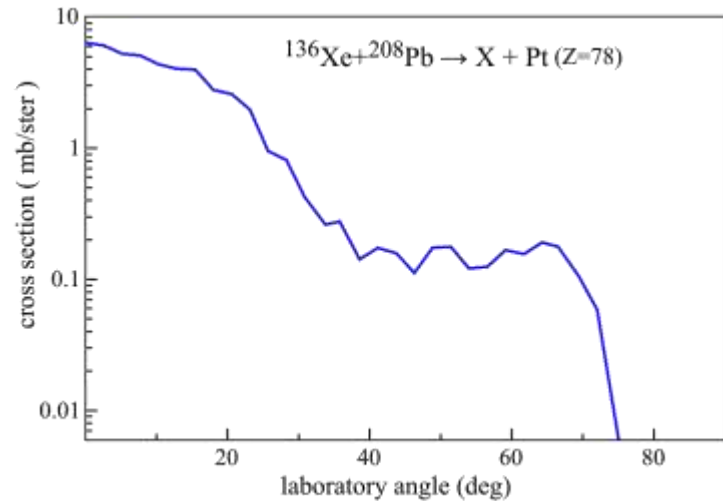
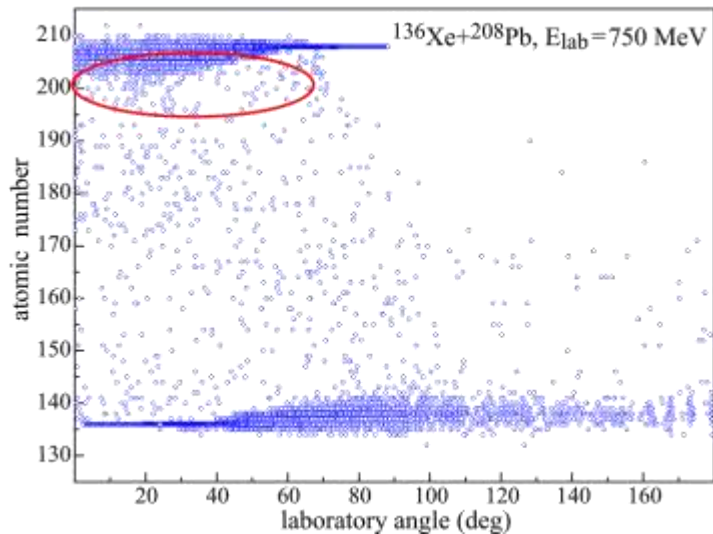
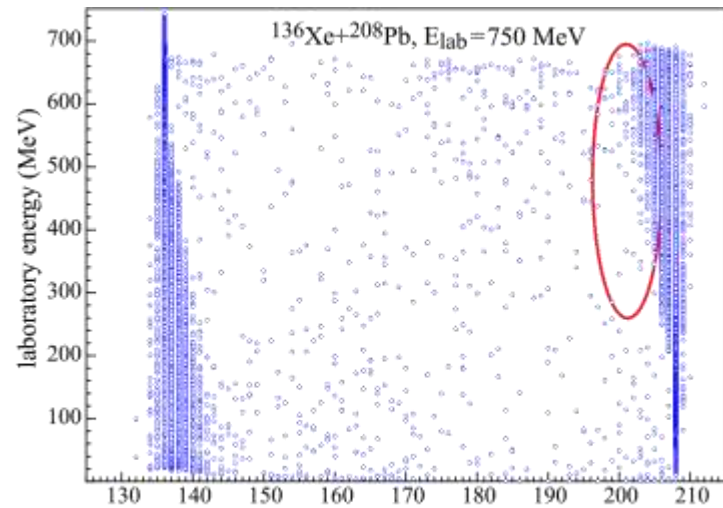
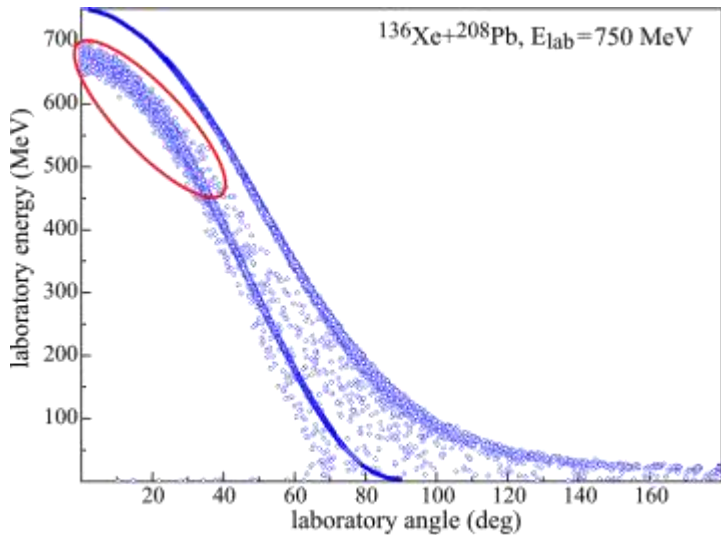
Several tens of new neutron-rich nuclides can be produced with cross section higher than one microbarn in the near-barrier collision of ^{136}Xe with ^{208}Pb



Production of heavy neutron-rich nuclei in transfer reactions with different stable and RI beams

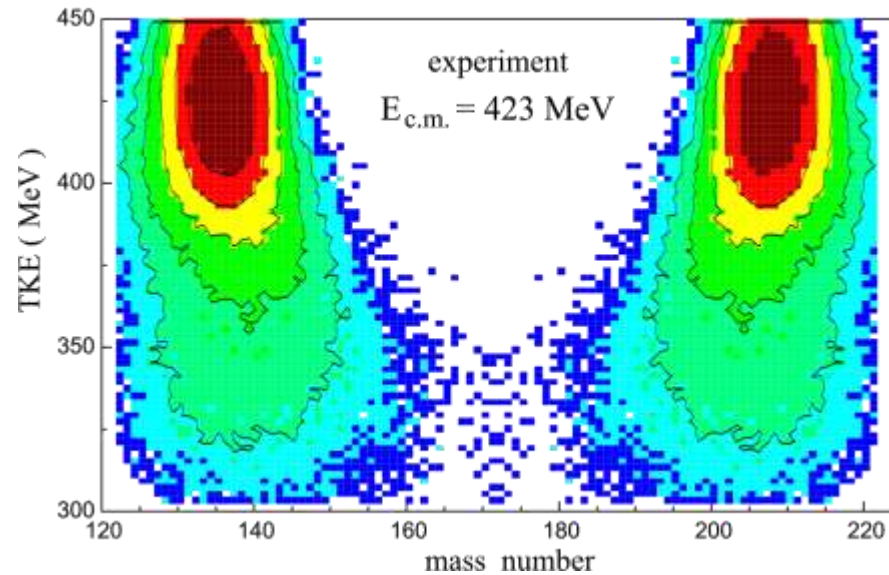
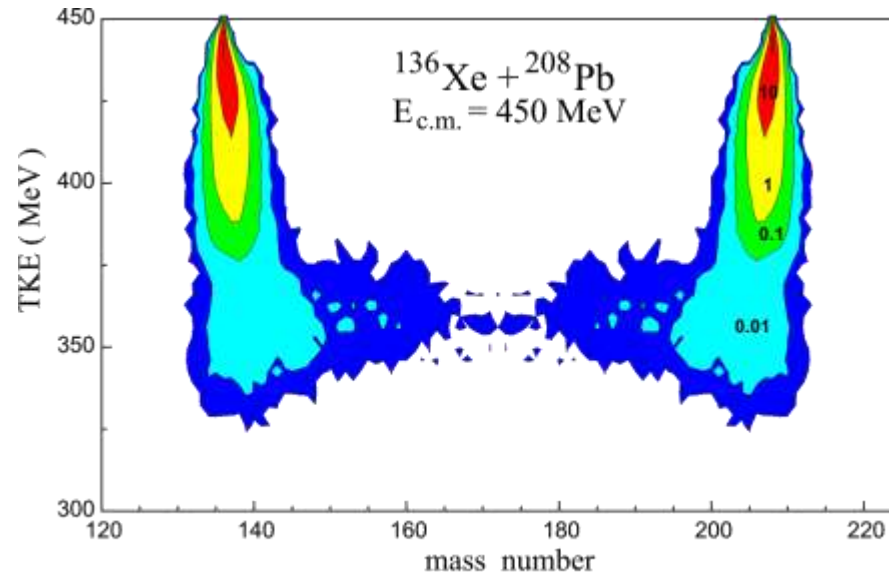


Simulation of typical experiment in the laboratory frame

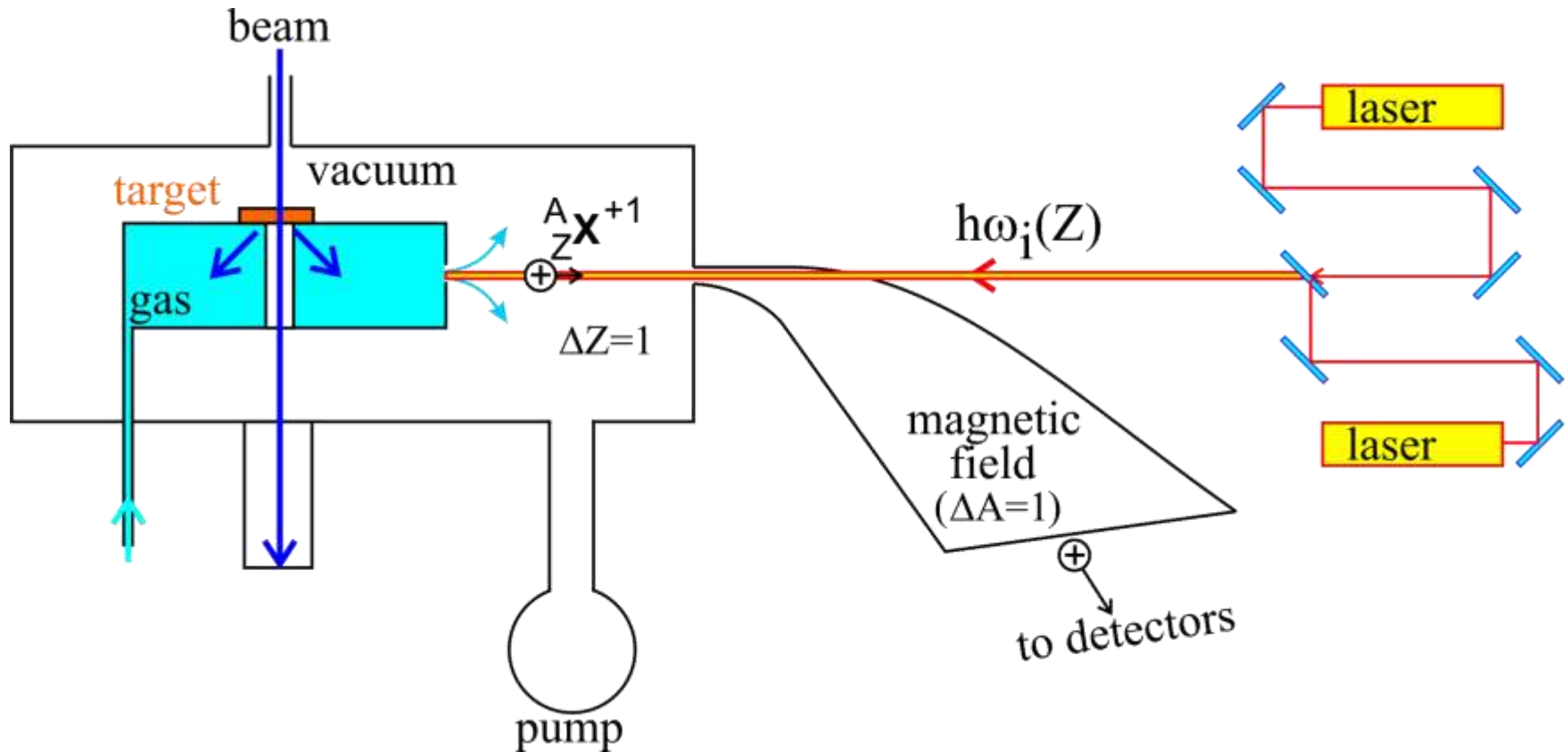


- (1) The yield of new neutron-rich isotopes is maximal at beam energy slightly above the Coulomb barrier
- (2) Desired reaction products are forward directed (no any grazing features)

Test experiment has demonstrated good agreement with our expectations



Schematic view of a new FLNR setup for resonance laser ionization of nuclear reaction products stopped in gas (IGISOL project)



Physics:
 study of new neutron-rich heavy nuclei produced in
 - multi-nucleon transfer reactions
 - quasi-fission process

Experiments started at:
 - FLNR (Dubna)
 - Jyvaskyla (Finland)
 Proposals submitted to:
 - GSI (Darmstadt)
 - Legnaro (Italy)

Summary

(nearest and distant experiments)

- Elements **119 and 120** may be synthesized next year in the Ti and/or Cr fusion reactions with cross sections of about 0.02 - 0.04 pb.
Perhaps they are the last SH elements with $T_{1/2} > 1 \mu\text{s}$?
- The **gap in SH mass area** ($Z = 105 - 116$) can be filled in fusion reactions of ^{48}Ca with lighter isotopes of actinides (^{239}Pu , ^{241}Am and ^{243}Cm). Cross sections are about 1 picobarn.
- Possible narrow **pathway to the island of stability** is found at last:
 $^{48}\text{Ca} + ^{250}\text{Cm} \rightarrow ^{295}116(\alpha) \rightarrow ^{291}114(\beta^+) \rightarrow ^{291}113(\beta^+) \rightarrow ^{291}112(T_{1/2} \sim 100 \text{ yrs}), \sigma \sim 0.8 \text{ pb} * 0.1 * 0.5$
- Multi-nucleon transfer reactions are to be used for synthesis of **neutron enriched long-living SH nuclei** close to the beta-stability line. **Uranium-like beams are needed !**
- Several tens of **new neutron-rich isotopes** of elements with $Z = 70 - 80$ (also those located along the closed neutron shell $N = 126$, astrophysical waiting point) can be produced in low-energy collisions of $^{136}\text{Xe} + ^{208}\text{Pb}$ with cross sections higher than one microbarn.
- A macroscopic amount of the long-living SH nuclei located at the island of stability may be produced with the use of **pulsed nuclear reactors** of the next generation (**factor 1000 is needed**).
- **Separators of a new kind are needed** for transfer reaction products with $Z > 70$ (**IGISOL ?**).
New detector equipment is also needed for studying new neutron rich beta-decaying nuclei.